



Quantum kinetic theory for chiral and spin transport in relativistic heavy ion collisions and core-collapse supernovae

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Quantum transport for massless fermions

Weyl fermions : chirality=helicity (no chirality mixing)

$$\overrightarrow{R}$$
 \overrightarrow{p} \overrightarrow{L}

$$\mathbf{J}_V = \mathbf{J}_R + \mathbf{J}_L$$

 $\mathbf{J}_5 = \mathbf{J}_R - \mathbf{J}_L$
(spin current)



Chiral anomaly: $\partial_{\mu}J^{\mu}_{R/L} = \pm \frac{\mathbf{E} \cdot \mathbf{B}}{4\pi^2} \implies \partial_{\mu}J^{\mu}_5 = \frac{\mathbf{E} \cdot \mathbf{B}}{2\pi^2}$

(Ohmic current)

Classical transport : $\mathbf{J}_V = \sigma_e \mathbf{E}$

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S. Adler, J. Bell, R. Jackiw, 69 K. Fujikawa, 79

 $A_{5}^{0} = \mu_{5}$

Anomalous transport (in chiral matter):

Chiral magnetic effect (CME): $\mathbf{J}_V = \frac{\mu_5}{2\pi^2} \mathbf{B}$ A. Vilenkin, PRD 20, 1807 (1979). PRD 22, 3080 (1980) K. Fukushima, D. Kharzeev, H. Warringa, PRD78, 074033 (2008) в D. E. Kharzeev, L. D. McLerran, H. J. Warringa, Nucl. Phys. A 803 (2008) 227-253





Chiral magnetic/vortical effect (CME/CVE)

Anomalous transport of Weyl fermions :

$$J^{\mu}_{R/L} = \pm \frac{\mu_{R/L}}{4\pi^2} B^{\mu} \pm \left(\frac{T^2}{12} + \frac{\mu^2_{R/L}}{4\pi^2}\right) \omega^{\mu}$$

A. Vilenkin, PRD 20, 1807 (1979). PRD 22, 3080 (1980)
K. Fukushima, D. Kharzeev, H. Warringa, PRD78, 074033 (2008)
D. T. Son, P. Surowka, PRL 103, 191601 (2009)
K. Landsteiner, E. Megias, F. Pena-Benitez, PRL107,021601(2011)

An "intuitive" picture for CVE : Barnett effect + spin enslavement

$$= \frac{1}{2} \nabla \times v$$

$$I_{R}$$

$$\vec{s} \neq \vec{r}$$

$$\vec{p} \neq \vec{L}$$

$$J_{L}$$

M. Matsuo et al, fphy.2015.00054

Keio University



Where to find anomalous transport in the real world?

 Nuclear physics : relativistic heavy ion collisions (strong B field & vorticity generated)

e.g. D. Kharzeev, et.al, Prog.Part.Nucl.Phys. 88, 1 (2016)

Condensed matter : Weyl semimetals

e.g. Q. Li, et.al., Nature Phys. 12 (2016) 550-554

Astrophysics : core-collapse supernovae

e.g. N. Yamamoto, Phys. Rev. D93, 065017 (2016)





Borisenko et al., Phys. Rev. Lett. 113, 027603 (2014)





- Outline
- Chiral and spin transport in relativistic heavy ion collisions (HIC)
- CME & spin polarization of hadrons
- Dynamical spin polarization : non-equilibrium spin transport
- Quantum kinetic theory (QKT)
- Spin-1/2 massless fermions
- Extensions (massive fermions, spin-1 massless bosons)
- Neutrino transport in core-collapse supernovae (CCSN)
- Chiral transport in CCSN
- Chiral magneto-hydrodynamics (CMHD) & inverse energy cascade
- Chiral radiation hydrodynamics (CRHD)
- Conclusions & outlook





- Global polarization of Λ hyperons :
- Statistical model/Wigner-function approach (in equilibrium):

$$\mathcal{P}^{\mu}(\mathbf{q}) = \frac{\int d\Sigma \cdot q J_{5}^{\mu}(q, X)}{2m \int d\Sigma \cdot \mathcal{N}(q, X)} \approx \frac{1}{8m} \epsilon^{\sigma \mu \nu \rho} q_{\sigma} \frac{\int d\Sigma \cdot q \omega_{\nu \rho} f_{q}^{(0)}(1 - f_{q}^{(0)})}{\int d\Sigma \cdot q f_{q}^{(0)}},$$
F. Becattini, et al., Ann. Phys. 338, 32 (2013)
R. Fang, et al., PRC 94, 024904 (2016)
$$\omega_{\nu \rho} = \frac{1}{2} \left(\partial_{\rho} (u_{\nu}/T) - \partial_{\nu} (u_{\rho}/T) \right)$$

thermal vorticity



Relativistic angular momentum

- Relativistic angular momentum for fermions : (review : E. Leader & C. Lorce, 13)
- Noether theorem :

• Canonical AM tensor :

$$\begin{split} M_{C}^{\lambda\mu\nu} &= M_{S}^{\lambda\mu\nu} + M_{O}^{\lambda\mu\nu}, \\ M_{S}^{\lambda\mu\nu} &= \frac{1}{2} \bar{\psi} \{\gamma^{\lambda}, \Sigma^{\mu\nu}\} \psi = \underbrace{-\frac{1}{2} \epsilon^{\lambda\mu\nu\rho} \bar{\psi} \gamma_{\rho} \gamma_{5} \psi}_{-\frac{1}{2} \epsilon^{\lambda\mu\nu\rho} \bar{\psi} \gamma_{5} \psi}_{-\frac{1}{2} \epsilon^{\lambda\mu\nu\rho} \bar{\psi}$$

 $\mathcal{L} = \bar{\psi} \left(\frac{i\hbar}{2} \gamma^{\mu} \overleftrightarrow{\partial}_{\mu} - m \right) \psi,$

 $\overleftrightarrow{\partial}_{\mu} = \overrightarrow{\partial}_{\mu} - \overleftarrow{\partial}_{\mu}$

• Canonical EM tensor: $T_C^{\mu\nu} = T^{\mu\nu} + T_A^{\mu\nu}, \ T^{\mu\nu} = \frac{i\hbar}{4}\bar{\psi}\gamma^{\{\mu}\overleftrightarrow{\partial}^{\nu\}}\psi, \ T_A^{\mu\nu} = \frac{i\hbar}{4}\bar{\psi}\gamma^{[\mu}\overleftrightarrow{\partial}^{\nu]}\psi.$

• EM & AM conservation : $\partial_{\mu}T_{C}^{\mu\nu} = 0,$ spin $-\frac{\hbar}{2}\epsilon^{\lambda\mu\nu\rho}\partial_{\lambda}J_{5\rho} + 2T_{A}^{\mu\nu} = 0$ $\partial_{\lambda}M_{C}^{\lambda\mu\nu} = 0.$



(same structure, opposite signs!)

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Dynamical evolution of the spin

- Disagreements also appear for transverse polarization
- (local) polarization may not be solely contributed by thermal vorticity

non-equilibrium effects may play a role
 W. Florkowski, et al., PRC100, 054907 (2019)
 H.-Z. Wu, et al., PRR 1, 033058 (2019)

Current theoretical studies :





Transport theories

- Kinetic theory : microscopic theory for quasi-particles in phase space
- Boltzmann (Vlasov) Eq. : $q^{\mu}\Delta_{\mu}f(q,X) = q^{\mu}\mathcal{C}_{\mu}[f], \quad \Delta_{\mu} = \partial_{\mu} + F_{\nu\mu}\frac{\partial}{\partial q^{\nu}}.$
- Physical quantities : $J^{\mu}(X) = \int \frac{d^3 q}{(2\pi)^3} \frac{q^{\mu}}{E_q} f(q, X), \quad T^{\mu\nu}(X) = \int \frac{d^3 q}{(2\pi)^3} \frac{q^{\mu} q^{\nu}}{E_q} f(q, X).$
- \clubsuit Validity : mean free path \gg de Broglie wavelength & mean free time \gg int. time



• Conservation laws : $\partial_{\mu}J^{\mu} = 0$, $\partial_{\mu}T^{\mu\nu} = F^{\nu\rho}J_{\rho}$.

✤ Near equilibrium : kinetic theory → hydrodynamics



Could we apply kinetic theory to HIC?

 Although QGP is strongly coupled, we may still learn some useful physics from the weakly coupled approach based on QCD.

(e.g. thermalization or hydrodynamization of HIC with kinetic theory) P. B. Arnold, G. D. Moore, L. G. Yaffe, JHEP 0301(2003) 030. A. Kurkela & Y. Zhu, PRL. 115, 182301 (2015)

 Primary goal : to construct a framework for tracking the spin transport of a probe fermion (s quark) traversing a weakly coupled medium (wQGP).

pros : microscopic theory, non-equilibrium,

Relativistic kinetic theory :

phase-space evolution

cons : weakly coupled, weak background fields

Quantum kinetic theory (QKT) : charge (energy-density) + spin dof

Cooper-Frye formula for spin : $\mathcal{P}^{\mu}(\mathbf{q}) = \frac{\int d\Sigma \cdot q J_5^{\mu}(q, X)}{2m \int d\Sigma \cdot \mathcal{N}(q, X)}$

F. Becattini, et al., Ann. Phys. 338, 32 (2013) R. Fang, et al., PRC 94, 024904 (2016)

Other approaches : microscopic models, spin hydro, holography etc.

e.g. J.-j. Zhang, et al., PRC 100, 064904 (2019) W. Florkowski, et al., PRC97, 041901 (2018) K. Hattori, et al., PLB 795, 100 (2019) A.D. Gallegos, U. Gürsoy, JHEP 11 (2020) 151



Chiral kinetic theory (CKT)

- Cara 1858 FORIO To study anomalous transport from kinetic theory :
 - Standard kinetic theory: $q^{\mu}\Delta_{\mu}f = q^{\mu}\mathcal{C}_{\mu}, \ \Delta_{\mu} = \partial_{\mu} + F_{\nu\mu}\frac{\partial}{\partial a^{\nu}} \implies \partial_{\mu}J^{\mu} = 0$
 - CKT: ? $\longrightarrow \partial_{\mu}J^{\mu} = \frac{\hbar}{4\pi^2} \mathbf{E} \cdot \mathbf{B}$ (for right-handed fermions)
 - Liouville's theorem : $\frac{d}{dt}f(x,\mathbf{p}) = (\partial_t + \mathbf{\dot{x}} \cdot \nabla_\mathbf{x} + \mathbf{\dot{p}} \cdot \nabla_\mathbf{p}) f = 0$
 - Quantum correction : classical action + Berry phase

D. T. Son & N. Yamamoto, PRL. 109, 181602 (2012)

• EOM: $\dot{\mathbf{x}} = \hat{\mathbf{p}} + \hbar \dot{\mathbf{p}} \times \mathbf{\Omega}_{\mathbf{p}}$, Berry curvature: $\mathbf{\Omega}_{p} = \frac{\lambda_{\mathrm{f}} p}{|p|^{3}}$, $\dot{\mathbf{p}} = \mathbf{E} + \dot{\mathbf{x}} \times \mathbf{B}.$ $\lambda_{\rm f} = \pm \frac{1}{2}.$

M. Stephanov & Y. Yin, PRL. 109, 162001 (2012)

 $(2|\mathbf{p}|^3)$

coupling

CKT (collisionless) :

$$\begin{bmatrix} (1 + \hbar \mathbf{B} \cdot \mathbf{\Omega}_{\mathbf{p}}) \partial_t + (\tilde{\mathbf{v}} + \hbar \mathbf{E} \times \mathbf{\Omega}_{\mathbf{p}} + \hbar (\tilde{\mathbf{v}} \cdot \mathbf{\Omega}_{\mathbf{p}}) \mathbf{B}) \cdot \nabla + \left(\tilde{\mathbf{E}} + \tilde{\mathbf{v}} \times \mathbf{B} + \hbar (\tilde{\mathbf{E}} \cdot \mathbf{B}) \mathbf{\Omega}_{\mathbf{p}}\right) \cdot \frac{\partial}{\partial \mathbf{p}} \end{bmatrix} f = 0$$
effective
$$\tilde{\mathbf{v}} = \partial \epsilon_{\mathbf{p}} / \partial \mathbf{p}$$

$$\epsilon_{\mathbf{p}} = |\mathbf{p}| - \frac{\hbar \mathbf{B} \cdot \mathbf{p}}{\hbar \mathbf{B} \cdot \mathbf{p}}$$
magnetic-moment

Field-theory derivation : Wigner functions + hbar expansion *

J.-W. Chen, et al., PRL. 110, 262301 (2013) D. T. Son & N. Yamamoto, PRD. 87, 085016 (2013)

Y. Hidaka, S. Pu, DY, PRD 95, 091901 (2017), PRD 97, 016004 (2018)

velocity/E field : $\tilde{\mathbf{E}} = \mathbf{E} - \partial \epsilon_{\mathbf{p}} / \partial \mathbf{x}$



Side jumps : spin enslavement + angular momentum (AM) conservation



★ frame-dep of f(p, x): modified Lorentz transformation.
J.-Y. Chen, et al., PRL. 113, 182302 (2014)
J.-Y. Chen, D. T. Son, and M. A. Stephanov,
PRL. 115, 021601 (2015).

quantum corrections on the collision term : self-energy gradients

Y. Hidaka, S. Pu, DY, PRD 95, 091901 (2017)

✤ In general, side jumps can also occur for free massless particles with spin



Wigner functions (WFs)

lesser (greater) propagators :

$$\tilde{S}_{\alpha\beta}^{>}(x,y) = \langle \psi_{\alpha}(x)U^{\dagger}(x,y)\bar{\psi}_{\beta}(y) \rangle$$

$$\tilde{S}_{\alpha\beta}^{<}(x,y) = \langle \bar{\psi}_{\beta}(y)U(y,x)\psi_{\alpha}(x) \rangle$$
gauge link
$$X = \frac{x+y}{2}, Y = x - y$$



review : J. Blaizot, E. Iancu, Phys.Rept. 359 (2002) 355-528

Wigner functions : $S_{\alpha\beta}^{<(>)}(q,X) = \int d^4Y e^{\frac{iq\cdot Y}{\hbar}} \tilde{S}_{\alpha\beta}^{<(>)}(x,y)$

- hbar expansion : $X \gg Y \rightarrow q_{\mu} \gg \partial_{\mu}$ (weak fields)
- Kadanoff-Baym (KB) equations :

$$(\not\!\!\!\!/ n - m)S^{<} + \gamma^{\mu}i\frac{\hbar}{2}\nabla_{\mu}S^{<} = \frac{i\hbar}{2}\left(\Sigma^{<} \star S^{>} - \Sigma^{>} \star S^{<}\right)$$

(systematic include collisions)

 $A \star B$

$$\begin{array}{l} \Delta_{\mu} = \partial_{\mu} + F_{\nu\mu}\partial/\partial q_{\nu} \\ \Delta_{\mu} = \partial_{\mu} + F_{\nu\mu}\partial/\partial q_{\nu} \\ \Pi^{\mu} = q^{\mu} + \mathcal{O}(\hbar^{2}) \\ = AB + \frac{i\hbar}{2} \{A, B\}_{\text{P.B.}} + \mathcal{O}(\hbar^{2}) \end{array}$$

$$\begin{array}{l} \mathbf{14} \end{array}$$

 $\nabla_{\mu} = \Delta_{\mu} + \mathcal{O}(\hbar^2).$



Quantum corrections for WFs

• R-handed WF up to $\mathcal{O}(\hbar)$: $S^{<}(q, X) = \bar{\sigma}^{\mu} \dot{S}_{\mu}(q, X)$,

$$\dot{S}^{<\mu}(q,X) = 2\pi \operatorname{sgn}(q\cdot n) \left(q^{\mu}\delta(q^2) f_q^{(n)} + \hbar\delta(q^2) S_{(n)}^{\mu\nu} \mathcal{D}_{\nu} f_q^{(n)} + \hbar\epsilon^{\mu\nu\alpha\beta} q_{\nu} F_{\alpha\beta} \frac{\partial\delta(q^2)}{2\partial q^2} f_q^{(n)} \right),$$

 $\mathcal{D}_{\beta}f_{q}^{(n)} = \Delta_{\beta}f_{q}^{(n)} - \Sigma_{\beta}^{<}(1 - f_{q}^{(n)}) + \Sigma_{\beta}^{>}f_{q}^{(n)}.$ side-jump term : magnetization current CVE or anomalous hall effects

 $\begin{array}{l} \text{spin tensor:} \ S^{\mu\nu}_{(n)} = \frac{\epsilon^{\mu\nu\alpha\beta}}{2(q\cdot n)} q_\alpha n_\beta \quad \text{J.-Y. Chen, et al., PRL. 113, 182302 (2014)} \end{array}$

> Frame vector n^{μ} : freedom for decomposition & a local observer we choose

The full WF has to be frame independent \$\impsycdots f_q^{(n)}\$ is frame dependent
 The modified frame transformation :

$$f_q^{(n')} = f_q^{(n)} + \frac{\hbar \epsilon^{\nu\mu\alpha\beta} q_\alpha n'_\beta n_\mu}{2(q \cdot n)(q \cdot n')} \mathcal{D}_\nu f_q^{(n)} \implies f^{(n')}(q, X) = f^{(n)}(q + \hbar \delta q, X + \hbar \delta X)$$
(side jumps)

modification on

the on-shell condition CME in equilibrium

> Y. Hidaka, S. Pu, DY, PRD 95, 091901 (2017), PRD 97, 016004 (2018)





CKT with collisions

• CKT with collisions $(\partial_{\rho}n^{\mu}=0)$:

Y. Hidaka, S. Pu, DY, PRD 95, 091901 (2017), PRD 97, 016004 (2018)

$$\delta\left(q^2 - \hbar \frac{B \cdot q}{q \cdot n}\right) \left\{ \left[q \cdot \Delta + \hbar \frac{S_{(n)}^{\mu\nu} E_{\mu}}{(q \cdot n)} \Delta_{\nu} + \hbar S_{(n)}^{\mu\nu} (\partial_{\mu} F_{\rho\nu}) \partial_{q}^{\rho} \right] f_{q}^{(n)} - \tilde{\mathcal{C}} \right\} = 0,$$

magnetic-moment coupling $(F^{\mu\nu} = 0 : \text{the quantum corrections only appear in collisions})$

Quantum corrections on the collision term :

$$\tilde{\mathcal{C}} = q \cdot \mathcal{C} + \hbar \frac{S_{(n)}^{\mu\nu} E_{\mu}}{(q \cdot n)} \mathcal{C}_{\nu} + \hbar S_{(n)}^{\alpha\beta} \underbrace{\left((1 - f_q^{(n)}) \Delta_{\alpha} \Sigma_{\beta}^{<} - f_q^{(n)} \Delta_{\alpha} \Sigma_{\beta}^{>} \right)}_{\text{induced by inhomoson}},$$

induced by inhomogeneity of the medium



• Solve $f_{R/L}$ and put them back to WFs $\implies A_{\mu}(q, X) = (\dot{S}_{\mu R}^{<} - \dot{S}_{\mu L}^{<})/2$



Anomalous (chiral) hydro from CKT

- Anomalous hydrodynamics (R-handed) : $\partial_{\mu}T^{\mu\nu} = F^{\nu\rho}J_{\rho}, \quad \partial_{\mu}J^{\mu} = \frac{\hbar}{4\pi^2}(\mathbf{E}\cdot\mathbf{B})$
- Constitutive equations:

$$T^{\mu\nu} = u^{\mu}u^{\nu}\epsilon - pP^{\mu\nu} + \Pi^{\mu\nu}_{\text{non}} + \Pi^{\mu\nu}_{\text{dis}}, \quad J^{\mu} = N_{0}u^{\mu} + v^{\mu}_{\text{non}} + v^{\mu}_{\text{dis}}, \quad P^{\mu\nu} = \eta^{\mu\nu} - u^{\mu}u^{\nu}.$$

$$\mathcal{O}(1) \qquad \mathcal{O}(\hbar) \quad \mathcal{O}(1) + \mathcal{O}(\hbar) \quad \mathcal{O}(1) \quad \mathcal{O}(\hbar) \quad \mathcal{O}(1) + \mathcal{O}(\hbar)$$

e.g. D. T. Son and P. Surowka, PRL 103, 191601 (2009)

depending on interactions

Y. Hidaka, S. Pu, DY, PRD 97, 016004 (2018)

$$T^{\mu\nu} = \int \frac{d^4q}{(2\pi)^4} \left[q^{\mu} \dot{S}^{<\nu} + q^{\nu} \dot{S}^{<\mu} \right], \quad J^{\mu} = 2 \int \frac{d^4q}{(2\pi)^4} \dot{S}^{<\mu}.$$

Transport coefficients from equilibrium WFs :

$$\implies v_{\rm non}^{\mu} = \hbar \sigma_B B^{\mu} + \hbar \sigma_{\omega} \omega^{\mu} \qquad \Pi_{\rm non}^{\mu\nu} = \hbar \xi_{\omega} \left(\omega^{\mu} u^{\nu} + \omega^{\nu} u^{\mu} \right) + \hbar \xi_B \left(B^{\mu} u^{\nu} + B^{\nu} u^{\mu} \right)$$

$$\sigma_{\omega} = \frac{T^2}{12} \left(1 + \frac{3\bar{\mu}^2}{\pi^2} \right), \ \sigma_B = \frac{\mu}{4\pi^2}, \ \xi_{\omega} = \frac{T^3}{6} \left(\bar{\mu} + \frac{\bar{\mu}^3}{\pi^2} \right), \ \xi_B = \frac{T^2}{24} \left(1 + \frac{3\bar{\mu}^2}{\pi^2} \right).$$
CVE
CME

(agree with those from different approaches, e.g. Son & Surowka, 09. K. Landsteiner, et.al. Lect. Notes, 13.)



$$\implies \partial_{\mu}s^{\mu}_{R\text{leq}} = 0$$

no entropy production

DY, PRD98, 076019 (2018)

See also D. T. Son and P. Surowka, PRL 103, 191601 (2009)

• Anti-sym EM tensor :
$$T_{A}^{\mu\nu} = \frac{\hbar}{2} N_{A} \left(\omega^{\mu} u^{\nu} - \omega^{\nu} u^{\mu} \right) \text{ from side-jumps}$$
DY, PRD98, 076019 (2018)
$$M_{\text{spin}}^{\lambda\mu\nu}(X) = \frac{\hbar}{2} \epsilon^{\lambda\mu\nu\rho} \left(N_{A} u_{\rho} + \hbar\sigma_{BA} B_{\rho} + \hbar\sigma_{\omega A} \omega_{\rho} \right) \text{ CSE \& CVE}$$

$$M_{\text{spin}}^{\lambda\mu\nu}(X) = \frac{\hbar}{2} \epsilon^{\lambda\mu\nu\rho} \left(N_{A} u_{\rho} + \hbar\sigma_{BA} B_{\rho} + \hbar\sigma_{\omega A} \omega_{\rho} \right) \text{ orbit}$$

 $-\frac{\pi}{2}\epsilon^{\lambda\mu\nu\rho}\partial_{\lambda}J_{5\rho} + 2T_{A}^{\mu\nu} = 0$

>
$$\mathcal{O}(\hbar)$$
: spin-orbit cancellation

Higher orders : we need higher-order WFs.



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Massive fermions

- Anomalous transport for massive fermions :
- Chiral symmetry is explicitly broken

> CSE & ACVE still exist :
$$J_5 = \sigma_{B5}B + \sigma_{\omega 5}\omega \xrightarrow[m \to \infty]{} J_5 \to 0$$

E. Gorbar, et al., PRD88, 025025 (2013) M. Buzzegoli, E. Grossi, F. Becattini, JHEP 10, 091 (2017) S. Lin and L. Yang, PRD98, 114022 (2018)

- QKT for massive fermions (e.g. for strange quarks)?
- > Chirality mixing : work in the vector/axial-vector bases.
- > Spin is no longer enslaved by chirality : a new dynamical dof
- > To track both vector/axial charges and spin polarization
- > Underlying quantum effects : chiral anomaly, spin-orbit int., etc.
- Axial kinetic theory (AKT) : a scalar + an axial-vector equations

K. Hattori, Y. Hidaka, DY, PRD100, 096011 (2019)
DY, K. Hattori, and Y. Hidaka, JHEP 20, 070 (2020)
N. Weickgenannt, et al., PRD 100 (2019), 056018.
J.-H. Gao & Z.-T. Liang, , PRD100 (2019), 056021.
Z. Wang, X. Guo, S. Shi, and P. Zhuang, PRD 100, 014015 (2019)
S. Li and H.-U. Yee, PRD100, 056022 (2019))
N. Weickgenannt, et al., arXiv:2005.01506
Z. Wang, X. Guo, P. Zhuang, arXiv:2009.10930 (see also my talk and the set of the set

(a modified relativistic BMT eq.)

V. Bargmann, L. Michel, and V.L. Telegdi, PRL 2, 435 (1959).

spin polarization

CME/CVE

$$\begin{array}{c} \square^{(n)} \mathcal{A}^{\mu} = \widehat{\mathcal{C}}^{\mu}_{cl} + \hbar \widehat{\mathcal{C}}^{(n)\mu}_{Q} \\ \text{spin diffusion} & \text{spin p} \end{array}$$

930 (see also my talk at Spin & Hydrodynamics online workshop by ect* on YouTube)

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Anomalous transport of photons

Photonic CVE :
$$j_{\text{CVE}}^{\pm} = \pm \frac{T^2}{6} \omega$$

N. Yamamoto, PRD96, 051902 (2017) See also : A. Avkhadiev and A. V. Sadofyev, PRD96, 045015 (2017) X.-G. Huang and A. V. Sadofyev, JHEP 03, 084 (2019)

Chern-Simons (CS) currents for R/L-handed photons

$$j_{\rm CVE}^{+\mu} - j_{\rm CVE}^{-\mu} \sim \mathcal{K}^{\mu} \equiv A_{\nu} \tilde{F}^{\mu\nu}$$

The gauge-inv. conserved quantity for Maxwell's equation :

 $\begin{array}{ll} \textbf{zilch:} & \partial_{\mu}Z^{\mu\nu\rho} = 0, \quad Z_{\mu\nu\rho} \equiv \frac{1}{2} \Big[F_{\mu}^{\ \alpha} \partial_{\rho} \tilde{F}_{\nu\alpha} - (\partial_{\rho}F_{\nu}^{\ \alpha}) \tilde{F}_{\mu\alpha} \Big] \\ & \text{D. Lipkin, J. Math. Phys. 5, 696 (1964)} \\ & \text{T. Kibble, J. Math. Phys. 6, 1022 (1965)} \\ \end{array}$

• Optical chirality :
$$Z_{000} = \frac{1}{2} (\mathbf{B} \cdot \dot{\mathbf{E}} - \mathbf{E} \cdot \dot{\mathbf{B}})$$

Y. Tang & A. Cohen, PRL. 104, 163901 (2010).

(asymmetry in the rates of excitation of chiral molecules)

Zilch vortical effect (ZVE) :

$$Z^{i} \equiv Z^{i00} = \frac{8\pi^{2}T^{4}\omega^{i}}{45} \quad \text{(zilch current)}$$

M. Chernodub, A. Cortijo, K. Landsteiner, PRD98, 065016 (2018)

C. Copetti, J. Fernandez-Pendas, PRD 98, 105008 (2018) (from Kubo formula)





WFs & QKT for photons

The full lesser propagator for photons (Coulomb gauge) : K. Hattori, Y. Hidaka, N. Yamamoto, DY, arXiv:2010.13368, JHEP XX (2021)

$$\begin{aligned} G_{\mu\nu}^{<} &\equiv G_{\mu\nu}^{\mathrm{R}<} + G_{\mu\nu}^{\mathrm{L}<} \\ &= 2\pi\delta(q^{2})\mathrm{sgn}(q \cdot n) \left[\begin{pmatrix} P \cdot \mathrm{even} \\ p_{\mu\nu}^{(n)} f_{\mathrm{V}} - \frac{\hbar q_{\perp(\mu} S_{\nu)\alpha}^{(n)} \partial^{\alpha}}{2(q \cdot n)^{2}} f_{\mathrm{A}} \end{pmatrix} - \mathrm{i} \begin{pmatrix} P \cdot \mathrm{odd} \\ S_{\mu\nu}^{(n)} f_{\mathrm{A}} + \frac{\hbar q_{\perp[\mu} \partial_{\perp\nu]}}{2(q \cdot n)^{2}} f_{\mathrm{V}} \end{pmatrix} \right], \\ f_{\mathrm{A}} &\equiv (f_{\mathrm{R}} - f_{\mathrm{L}})/2 \text{ and } f_{\mathrm{V}} \equiv (f_{\mathrm{R}} + f_{\mathrm{L}})/2 \\ & \mathsf{for CS \& Zilch currents} \\ & \mathsf{L.O. polarization tensors} : P_{\mu\nu}^{(n)} = n_{\mu}n_{\nu} - \eta_{\mu\nu} - \hat{q}_{\perp\mu}\hat{q}_{\perp\nu}, \quad S_{\mu\nu}^{(n)} = \frac{\epsilon_{\mu\nu\alpha\beta}q^{\alpha}n_{\beta}}{q \cdot n} \\ & \mathsf{Spin tensor for photons} \\ & \mathsf{CVE \& ZVE} : \quad \mathcal{K}_{\mathrm{eq}}^{\mu} = \frac{2}{9}T^{2}\omega^{\mu} \quad \& \quad Z_{\mathrm{eq}}^{\alpha}(X) = \frac{8\pi^{2}}{45\hbar^{2}}T^{4}\omega^{\alpha} \quad \overset{\mathrm{see also}}{X \cdot \mathrm{G. Huang, et al., JHEP 10, 117 (2020)} \\ & \mathsf{Effective QKT for photons (gluons): } q \cdot \partial f_{\mathrm{A}}^{(\gamma)} = \mathcal{C}_{\mathrm{A}}^{(\gamma)} \\ & + \\ \Box^{(n)}\mathcal{A}^{\mu} = \hat{\mathcal{C}}_{\mathrm{cl}}^{\mu} + \hbar \hat{\mathcal{C}}_{\mathrm{O}}^{(n)\mu} \\ & \mathsf{Coupled spin evolution in wQGP} \end{aligned}$$



Applications in astrophysics

 Core-collapse supernovae (CCSN) could be ideal for studying the quantum transport of chiral matter. N. Yamamoto, PRD93, 065017 (2016)

Y. Masada, K. Kotake, T. Takiwaki, and N. Yamamoto, PRD98, 083018 (2018)

- Intrinsic chiral matter created in nature : weak interaction
- Neutrino/electron transport in supernovae :







Inverse energy cascade

Normal matter : direct cascade



https://doi.org/10.1515/htmp-2016-0043

 Chiral matter : inverse cascade (with chiral anomaly)

(From the small scale to the large scale) could favor supernova explosions chiral magneto-hydrodynamic (ChMHD) simulations :

Y. Masada, K. Kotake, T. Takiwaki, and N. Yamamoto, PRD 98 (2018) 8, 083018

(From the large scale to the small scale) standard 3-d magneto-hydrodynamic (MHD) simulations : disfavor robust supernova explosions

(as opposed to the 2-d case)





Chirality transfer & chiral plasma instability

Chirality (helicity) transfer characterized by helicity conservation :

For Weyl fermions :
$$\partial_{\mu} j^{\mu} = -CE^{\mu}B_{\mu}$$

$$\implies \frac{d}{dt} \int d^{3}x \left(j^{0} + \frac{C}{2} \mathbf{A} \cdot \mathbf{B} \right) = 0, \quad j^{0} = n + \underbrace{\xi v \cdot \omega + \xi_{B} v \cdot \mathbf{B}}_{\text{from CVE & CME}}.$$

$$\implies \text{ conservation law : } \frac{d}{dt}Q_{\text{tot}} = 0, \quad Q_{\text{tot}} \equiv Q_{\text{chi}} + \frac{C}{2}Q_{\text{mag}} + \xi Q_{\text{flu}} + \xi_{B}Q_{\text{mix}},$$

$$Q_{\text{chi}} = \int d^{3}x \ n, \quad Q_{\text{mag}} = \int d^{3}x \ \mathbf{A} \cdot \mathbf{B}, \quad Q_{\text{flu}} = \int d^{3}x \ \mathbf{v} \cdot \boldsymbol{\omega}, \quad Q_{\text{mix}} = \int d^{3}x \ \mathbf{v} \cdot \mathbf{B}.$$

Chiral plasma instability (CPI): Y. Akamatsu and N. Yamamoto, PRL 111, 052002 (2013).

Unstable mode of magnetic fields : $\delta m{B} \propto \exp(i m{k} \cdot m{x} + \sigma t)$

$$\sigma = \eta \xi_B k - \eta k^2$$
 (small k)

 \propto initial chirality imbalance



Potential applications in astrophysics

- Neutrino transport in core-collapse supernovae
- Weak int. yields intrinsic chiral imbalance
- Neutrinos are out of equilibrium away from the core.



- Near the core : ChMHD (e, N, v)
- Away from the core : ChMHD (e, N) +CKT (ν) \implies chiral radiation hydrodynamics

$$\nabla_{\mu}T_{\rm rad}^{\mu\nu} + \nabla_{\mu}T_{\rm mat}^{\mu\nu} = 0$$

(neutrinos)

(electrons, nucleons : equilibrium, non-relativistic)





Chiral radiation transport equation

- CKT for neutrinos : $q \cdot Df_{Lq}^{(\xi)} = (1 f_{Lq}^{(\xi)})\Gamma_{(\xi)q}^{<} f_{Lq}^{(\xi)}\Gamma_{(\xi)q}^{>}, \quad n^{\mu} = \xi^{\mu} = (1, 0).$ (matter (n, p, e) in equilibrium) N. Yamamoto & DY, APJ 895 (2020), 1
- Em/ab rates : $\Gamma_{(\xi)q}^{\leq} \approx \Gamma_q^{(0)\leq} + \hbar \Gamma_q^{(\omega)\leq} (q \cdot \omega) + \hbar \Gamma_q^{(B)\leq} (q \cdot B)$

 $M_{\rm n} \approx M_{\rm p} \approx M$ NR approx., $\omega^{\mu} \equiv \frac{1}{2} \epsilon^{\mu}$ small-momentum transf.

$$\nu_{\mathrm{L}}^{\mathrm{e}}(q) + \mathrm{n}(k) \rightleftharpoons \mathrm{e}_{\mathrm{L}}(q') + \mathrm{p}(k')$$

 $\omega^{\mu} \equiv \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} u_{\nu} (\partial_{\alpha} u_{\beta})$ vorticity & magnetic field corrections :

- \succ incorporating $v \cdot \omega \& v \cdot B$ terms
- breaking spherical symmetry & axisymmetry

analytic expressions : $\Gamma_q^{(0)>} \approx \frac{G_{\rm F}^2}{\pi} \left(g_{\rm V}^2 + 3g_{\rm A}^2\right) E_{\nu}^3 \left(1 - f_q^{(\rm e)}\right) \left(1 - \frac{3E_{\nu}}{M_{\rm N}}\right) \frac{n_{\rm p} - n_{\rm n}}{1 - {\rm e}^{\beta(\mu_{\rm n} - \mu_{\rm p})}}$

$$\Gamma_q^{(B)>} \approx \frac{G_{\rm F}^2}{2\pi M_{\rm N}} \left(g_{\rm V}^2 + 3g_{\rm A}^2\right) E_{\nu} \left(1 - f_q^{\rm (e)}\right) \left(1 - \frac{8E_{\nu}}{3M_{\rm N}}\right) \frac{n_{\rm p} - n_{\rm n}}{1 - {\rm e}^{\beta(\mu_{\rm n} - \mu_{\rm p})}}$$

$$\Gamma_q^{(\omega)>} \approx \frac{G_{\rm F}^2}{2\pi} \left(g_{\rm V}^2 + 3g_{\rm A}^2 \right) E_{\nu} (1 - f_q^{\rm (e)}) \left(2 + \beta E_{\nu} f_q^{\rm (e)} \right) \frac{n_{\rm p} - n_{\rm n}}{1 - {\rm e}^{\beta(\mu_{\rm n} - \mu_{\rm p})}}$$

 $\Gamma_q^{(0)>}$: S. Reddy, M. Prakash, J. M. Lattimer, PRD58:013009,1998





Conclusions & outlook

- The QKT for fermions from WFs provides a useful theoretical framework to track non-equilibrium spin and chiral transport in phase space.
- To delineate the dynamical spin polarization of a strange quark traversing wQGP, the QKT of massive fermions is constructed.
- Similar to massless fermions, photons & gluons may have exotic transport properties under rotation. The QKT of photons/gluons is derived, which can be coupled to the QKT of fermions for the application to wQGP.
- Details for the collision term in wQCD has to be worked out.
- Chiral transport of leptons could have potential impacts on the evolution of CCSN.
- Qualitatively, the chirality imbalance could lead to inverse energy cascade and facilitate the explosions.
- The chiral transport equation of neutrinos has been constructed from CKT, which should be further applied to the simulations of CCSN.
- Further simplification or extension of ChRHD will be needed.
- Other applications of QKT to e.g. condensed-matter physics?



Thank you!



WFs in thermal equilibrium



Constitutive equations :

 $T^{\mu\nu} = u^{\mu}u^{\nu}\epsilon - pP^{\mu\nu} + \Pi^{\mu\nu}_{non} + \Pi^{\mu\nu}_{dis}, \quad J^{\mu} = N_{0}u^{\mu} + v^{\mu}_{non} + v^{\mu}_{dis}, \quad P^{\mu\nu} = \eta^{\mu\nu} - u^{\mu}u^{\nu}.$ $\mathcal{O}(1) \qquad \mathcal{O}(\hbar) \quad \mathcal{O}(1) + \mathcal{O}(\hbar) \quad \mathcal{O}(1) \quad \mathcal{O}(\hbar) \quad \mathcal{O}(1) + \mathcal{O}(\hbar)$ D. T. Son and P. Surowka, PRL 103, 191601 (2009) equilibrium : CME/CVE 29



■ Massive fermions : Reducing redundant dof : replacing S, P, and $S^{\mu\nu}$ in terms of \mathcal{V}^{μ} and \mathcal{A}^{μ} (10 eqs. \rightarrow 6 master eqs.).

e.g.
$$m\mathcal{P} = -\frac{\hbar}{2}\nabla_{\mu}\mathcal{A}^{\mu} \implies \partial_{\mu}J_{5}^{\mu} = \frac{\hbar\mathbf{E}\cdot\mathbf{B}}{2\pi^{2}} + 2im\langle\bar{\psi}\gamma_{5}\psi\rangle$$

Massless limit:
$$\dot{S}^{<} = \begin{pmatrix} 0 & \sigma^{\mu} (\mathcal{V}_{\mu} - \mathcal{A}_{\mu}) \\ \bar{\sigma}^{\mu} (\mathcal{V}_{\mu} + \mathcal{A}_{\mu}) & 0 \end{pmatrix} = \begin{pmatrix} 0 & \sigma^{\mu} \dot{S}_{L\mu}^{<} \\ \bar{\sigma}^{\mu} \dot{S}_{R\mu}^{<} & 0 \end{pmatrix}$$
LO WFs: $\dot{S}_{R/L\mu}^{<} = 2\pi \bar{\epsilon}(q_0) q_{\mu} f_{R/L}(q, X), \quad f_{V/A} \equiv \frac{(f_R \pm f_L)}{2}.$



Axial kinetic theory

- QKT : modified Boltzmann eqs. tracking charge/spin evolution of a quasiparticle in phase space with quantum corrections.
- > Vector/axial-vector $\mathcal{V}^{\mu}/\mathcal{A}^{\mu}$: $f_{V/A}(q,X)$ & $a^{\mu}(q,X)$ spin 4-vector (2 dof)

$$\xrightarrow{n=0} a^{\mu} = q^{\mu}, \quad f_{V/A} \equiv \frac{(f_{R} \pm f_{L})}{2}$$

 $\hbar \hat{\mathcal{C}}_{\widehat{}}^{(n)\mu}$

- Axial kinetic theory : scalar/axial-vector kinetic eqs. (SKE/AKE)
- > To include collisions : $\mathcal{V}^{\mu} \sim \mathcal{O}(\hbar^0)$ and $\mathcal{A}^{\mu} \sim \mathcal{O}(\hbar)$

Local equilibrium :

with quantum corrections

> Effective QKT for spin :

DY, K. Hattori, and Y. Hidaka, JHEP 20, 070 (2020)

(see also S. Li and H.-U. Yee, PRD100, 056022 (2019))

The formalism is derived yet the details need to be worked out.

 $\Box^{(n)} \mathcal{A}^{\mu}$

E.g. Vorticity-induced spin polarization in Nambu–Jona-Lasinio (NJL) model:

detailed balance ⇒

$$\mathcal{A}^{\rm LE}_{\mu}(p) = \mathcal{A}^{\rm LE(0)}_{\mu}(p) + \hbar \mathcal{A}^{\rm LE(1)}_{\mu}(p) = -\frac{\hbar}{(2\pi)^3 2E_p} \epsilon_{\mu\nu\sigma\lambda} p^{\nu} \nabla^{\sigma} \beta^{\lambda} f'_{V,\rm LE}(X,p)$$

(matches the result in global equilibrium)

Z. Wang, X. Guo, P. Zhuang, arXiv:2009.10930



 $\nu_{\mathrm{L}}^{\ell}(q) + \mathrm{N}(k) \rightleftharpoons \nu_{\mathrm{L}}^{\ell}(q') + \mathrm{N}(k')$