

Toward QCD-based description of dense baryonic matter

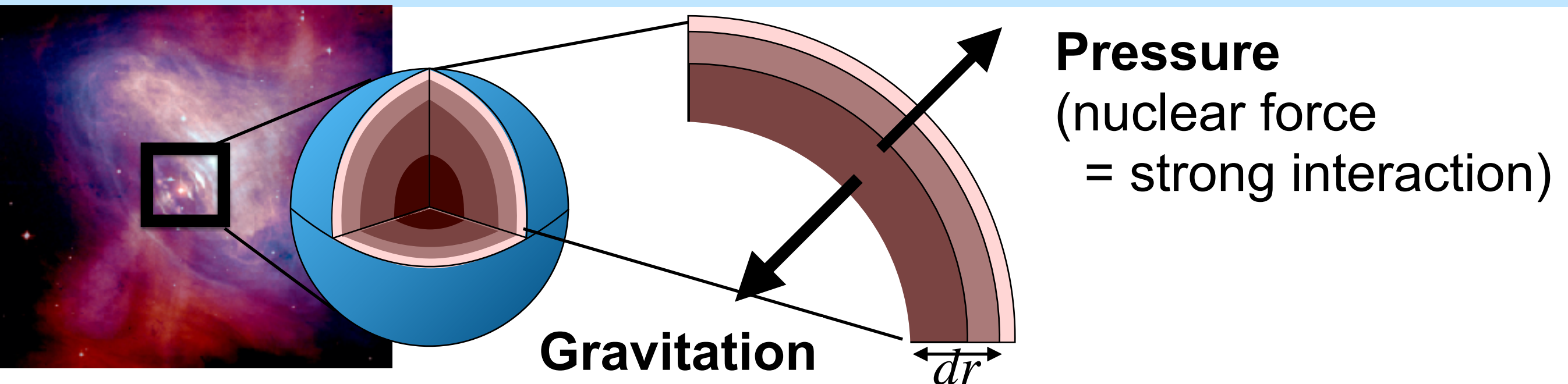
Yuki Fujimoto

(The University of Tokyo)

Reference:

Y. Fujimoto, K. Fukushima, “Equation of state of cold and dense QCD matter in resummed perturbation theory,” arXiv:2011.10891.

Neutron star structure



Hydrostatic equilibrium (pressure = gravitation)

$$\frac{dp(r)}{dr} = -G \frac{m(r)\epsilon(r)}{r^2} \times \underbrace{\left(1 + \frac{p}{\epsilon}\right) \left(1 + \frac{4\pi r^3 p}{m}\right) \left(1 - \frac{2Gm}{r}\right)^{-1}}_{\text{General relativistic correction}}$$

Tolman (1939)
Oppenheimer, Volkoff (1939)

← **TOV equation**

$$m(r) = \int_0^r dr 4\pi r^2 \epsilon(r)$$

Unknown variables:

$p(r)$, $m(r)$ and $\epsilon(r)$

One condition
missing!

Equation of State (EoS)

$$p = p(\epsilon)$$

Neutron star structure

TOV equation:

Tolman (1939)
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Equation of State (EoS): $p = p(\epsilon)$

Initial value problem

Initial:

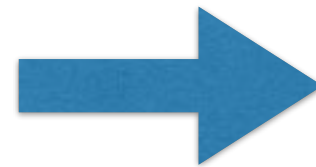
$$r = 0$$

$$\epsilon(r = 0) = \epsilon_c$$

free parameter

$$p(r = 0) = p(\epsilon_c) = p_c$$

$$m(r = 0) = 0$$



Final:

$$r = \mathbf{R} \quad \mathbf{M-R} \text{ relation}$$

$$\epsilon(r = R) = 0$$

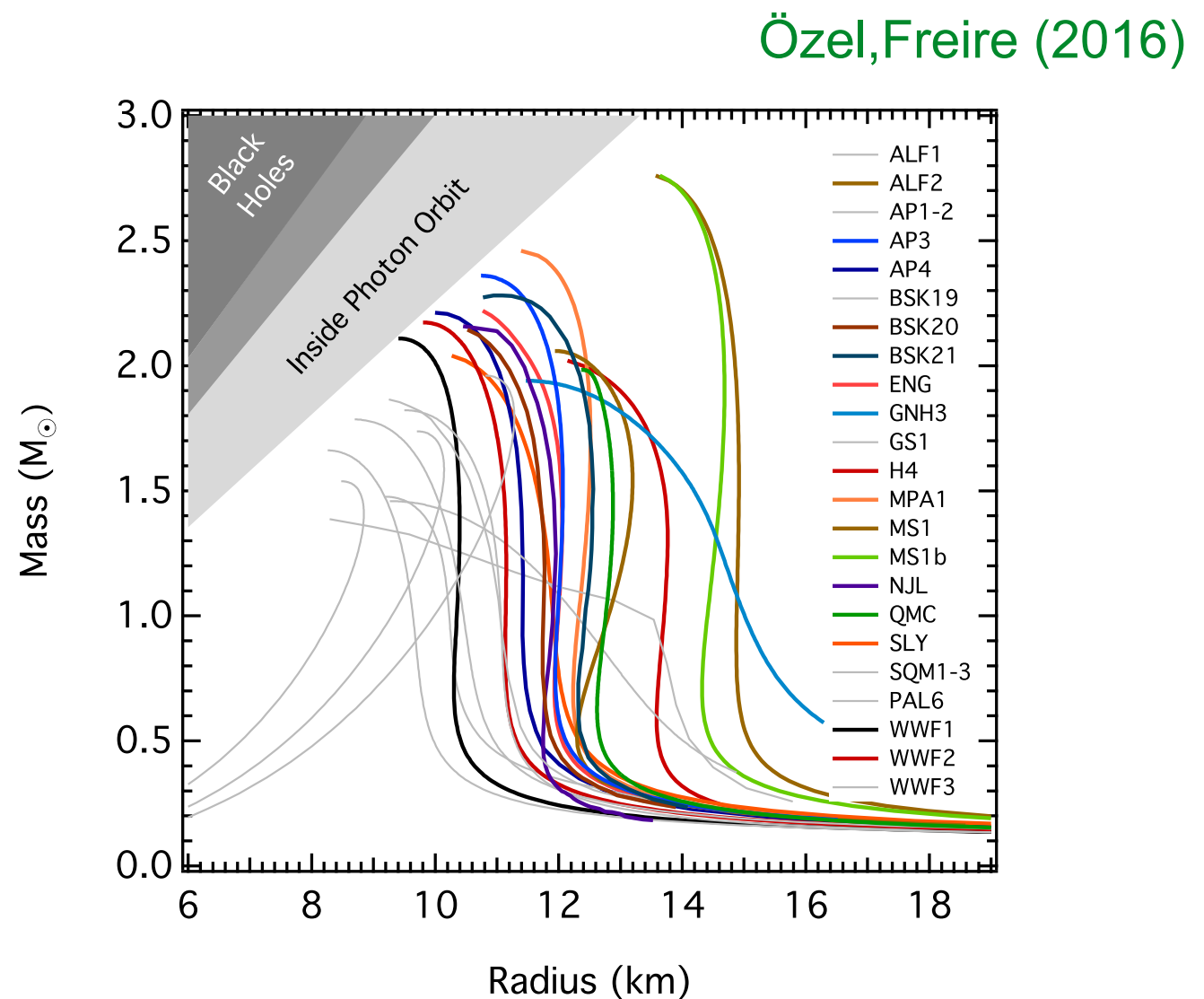
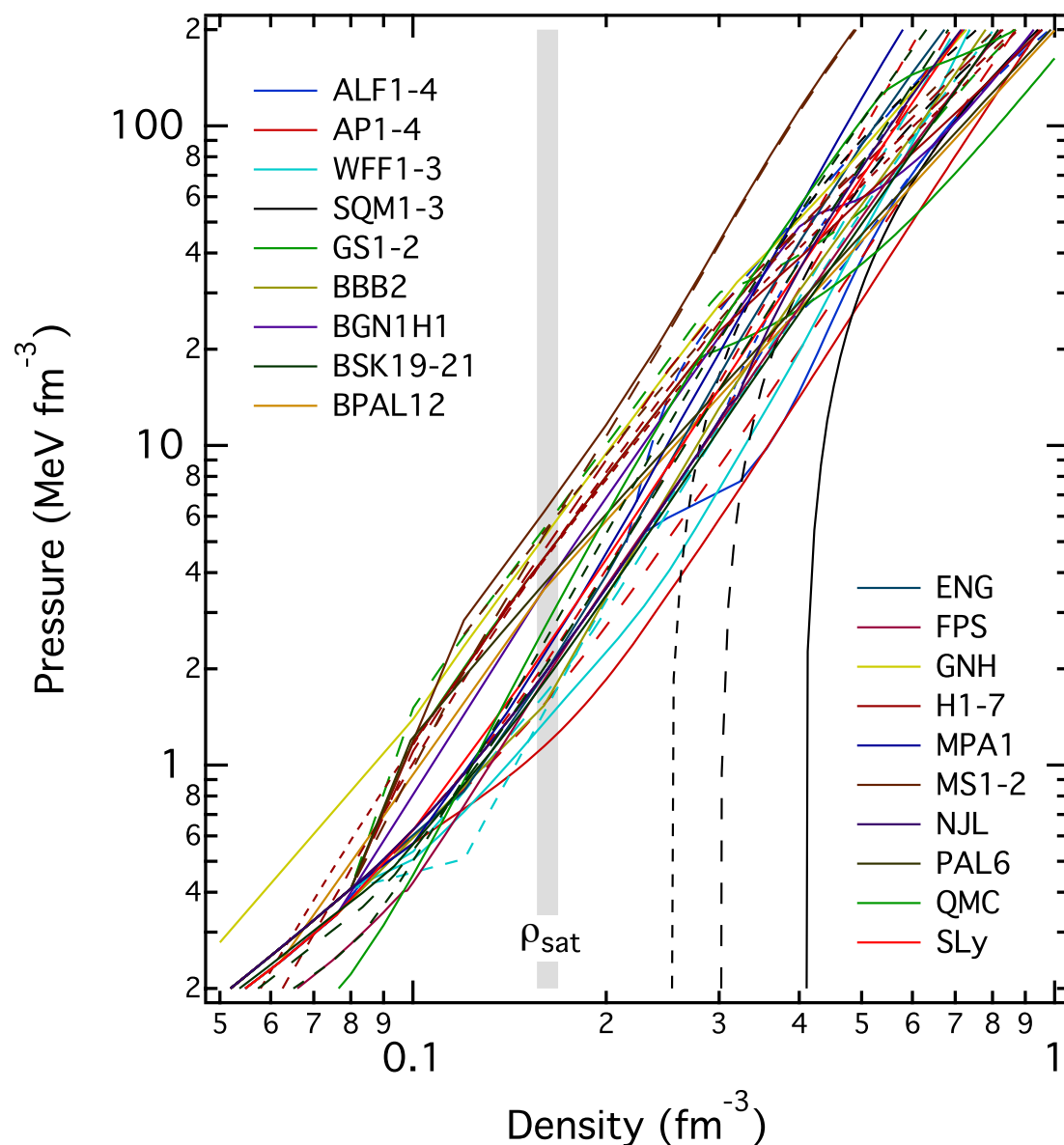
$$p(r = R) = 0$$

$$m(r = R) = \mathbf{M}$$

Equation of state (EoS)

EoS: pressure function $p(n_B)$, $p(\varepsilon)$, or $p(\mu_B)$

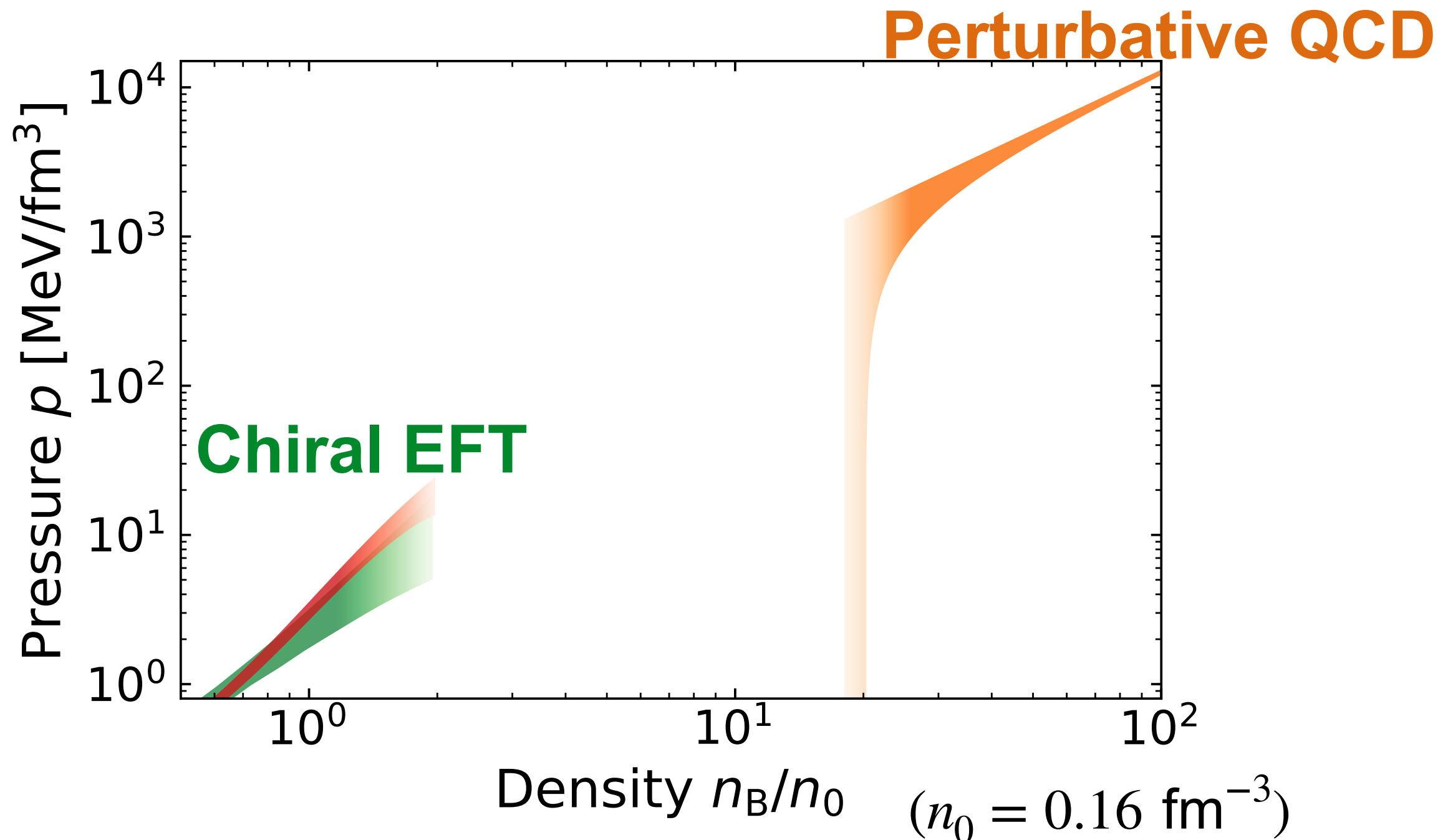
(n_B : baryon density, ε : energy density, μ_B : chemical potential)



“Spaghetti diagram”, a little messy... 4

A QCD physicist's view on the NS EoS

Constraints from QCD point of view:

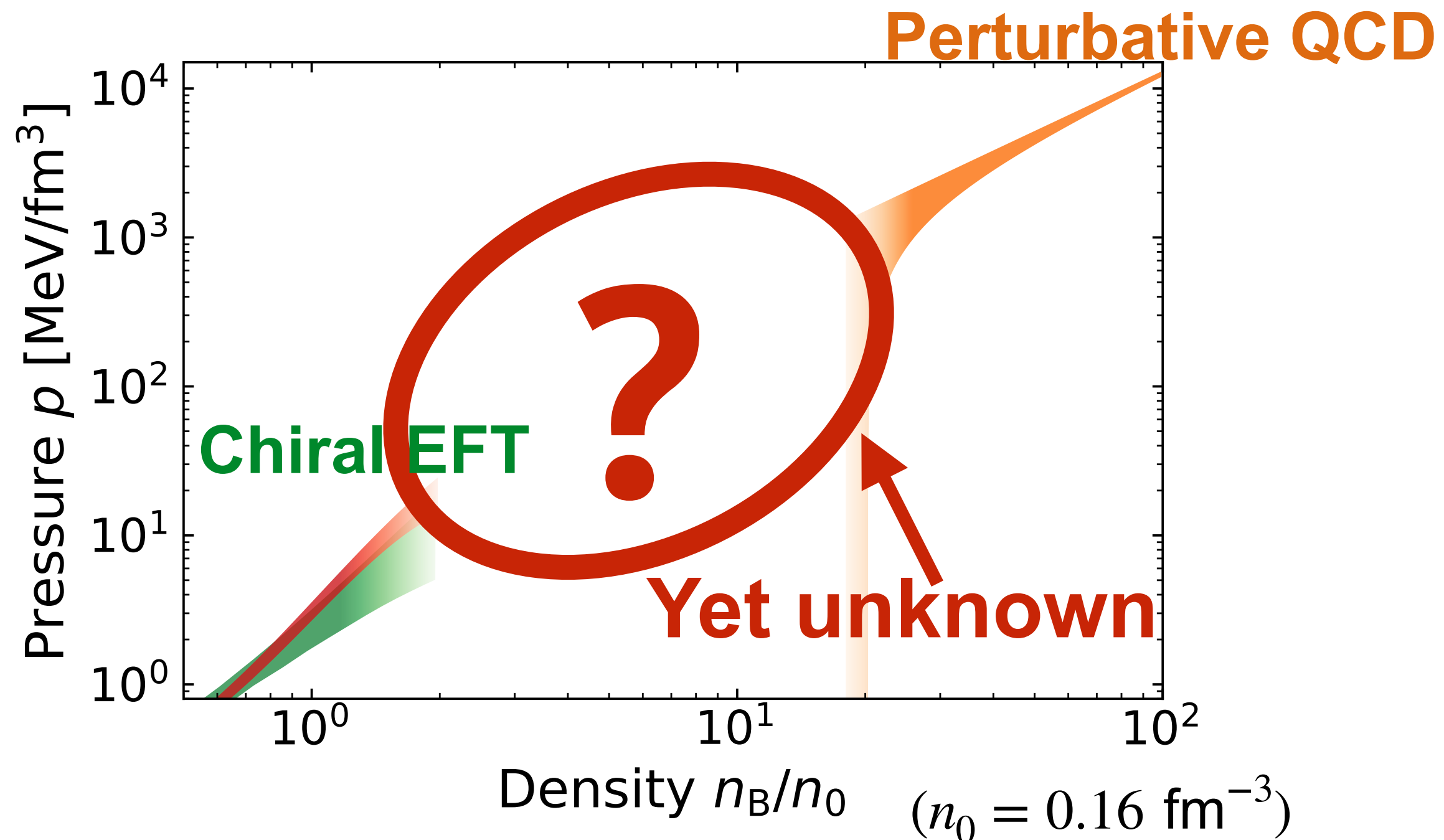


ChEFT: Tews, Carlson, Gandolfi, Reddy (2018);
Drischler, Furnstahl, Melendez, Phillips (2020)

pQCD: Freedman, McLerran (1978); Baluni (1979);
Kurkela, Romatschke, Vuorinen, Gorda, Sappi (2009-) 5

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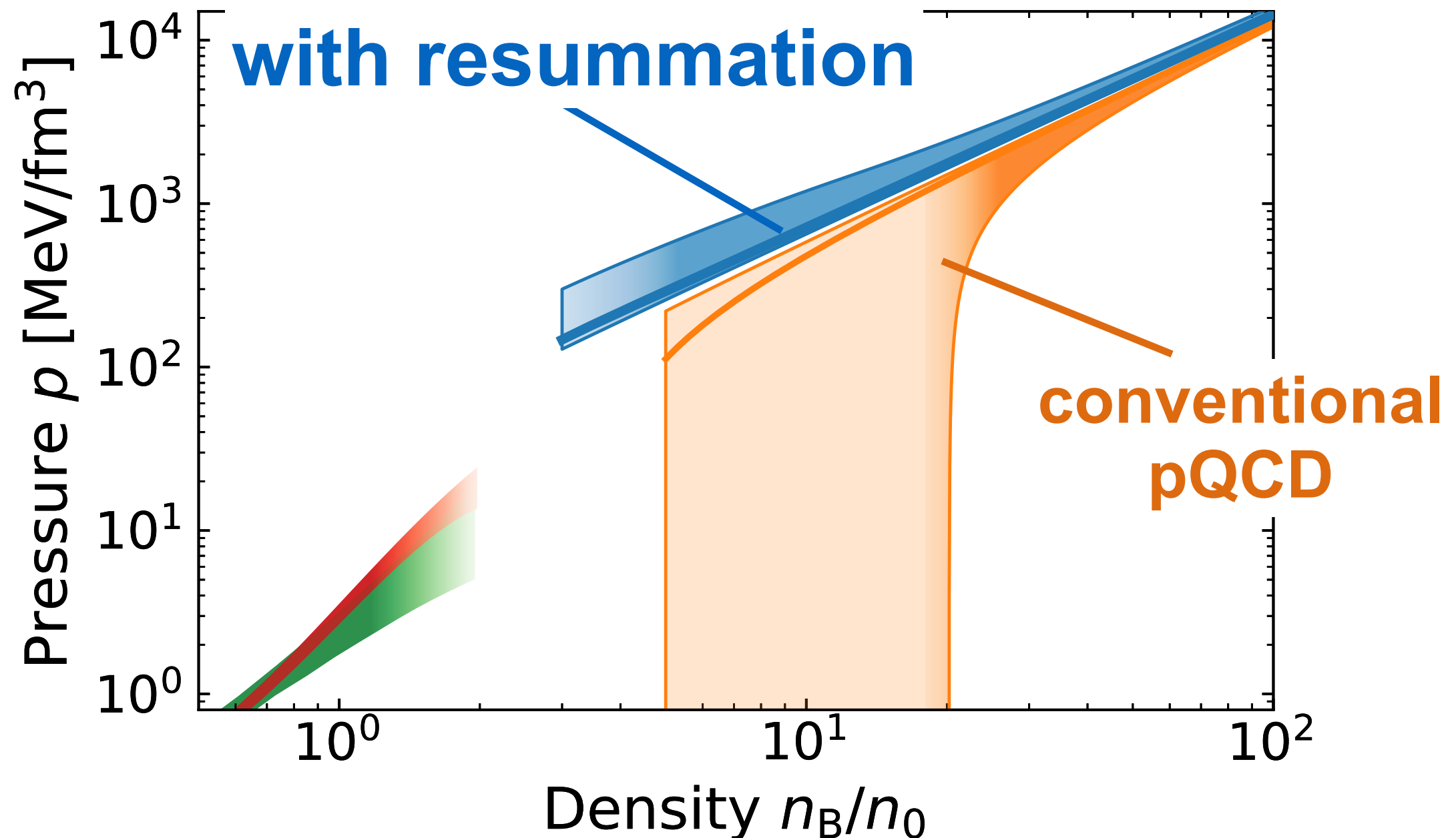
Our result in a nutshell

$$\alpha_s(\bar{\Lambda}) = \frac{4\pi}{\beta_0 \log(\bar{\Lambda}^2/\Lambda_{\overline{\text{MS}}}^2)} \left[1 - \frac{2\beta_1}{\beta_0^2} \frac{\log^2(\bar{\Lambda}^2/\Lambda_{\overline{\text{MS}}}^2)}{\log(\bar{\Lambda}^2/\Lambda_{\overline{\text{MS}}}^2)} \right]$$

uncertainty: originate from
scale variation

$$\bar{\Lambda} = \mu, 2\mu, 4\mu$$

**Our result
with resummation**



Our result in a nutshell

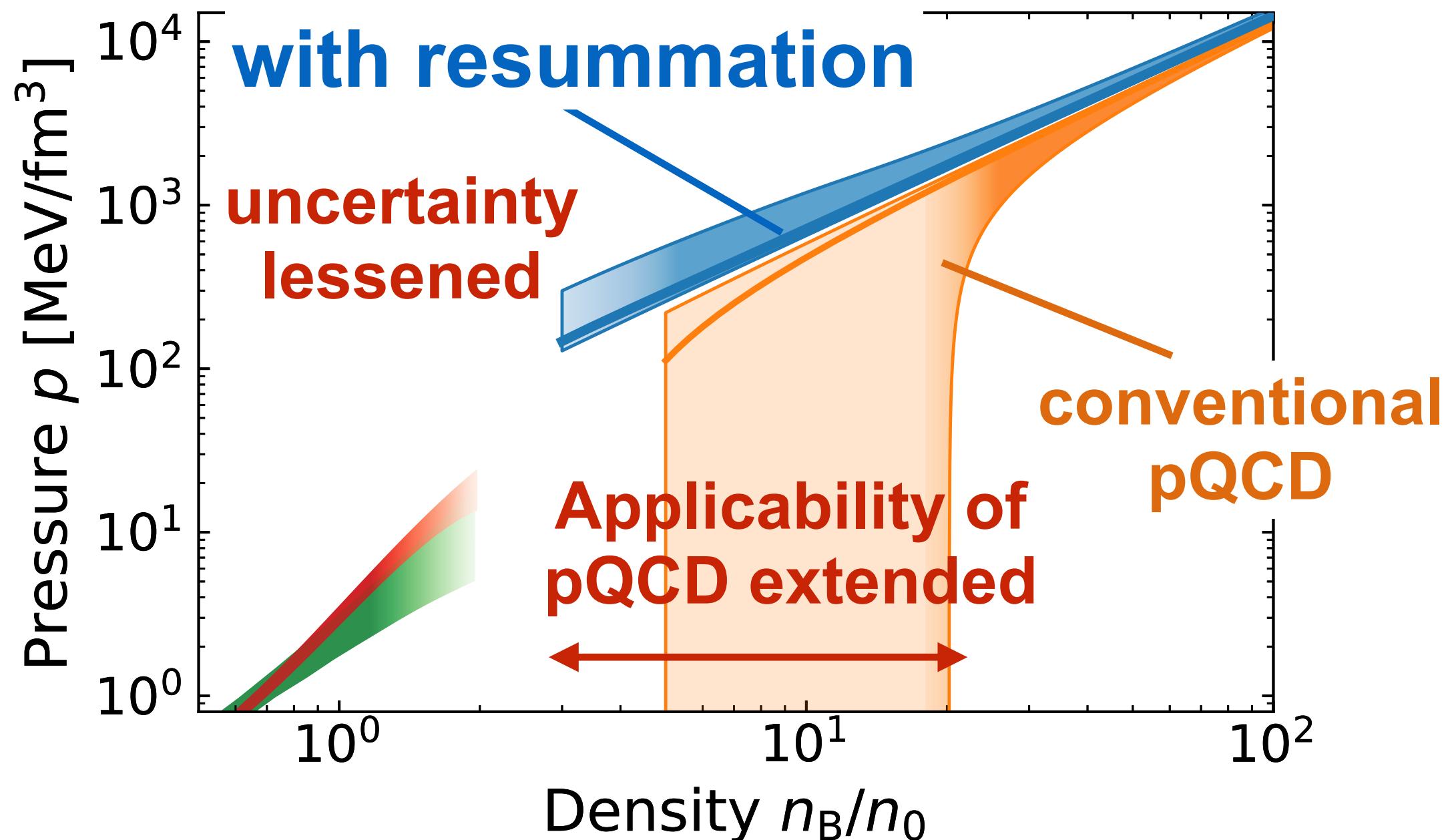
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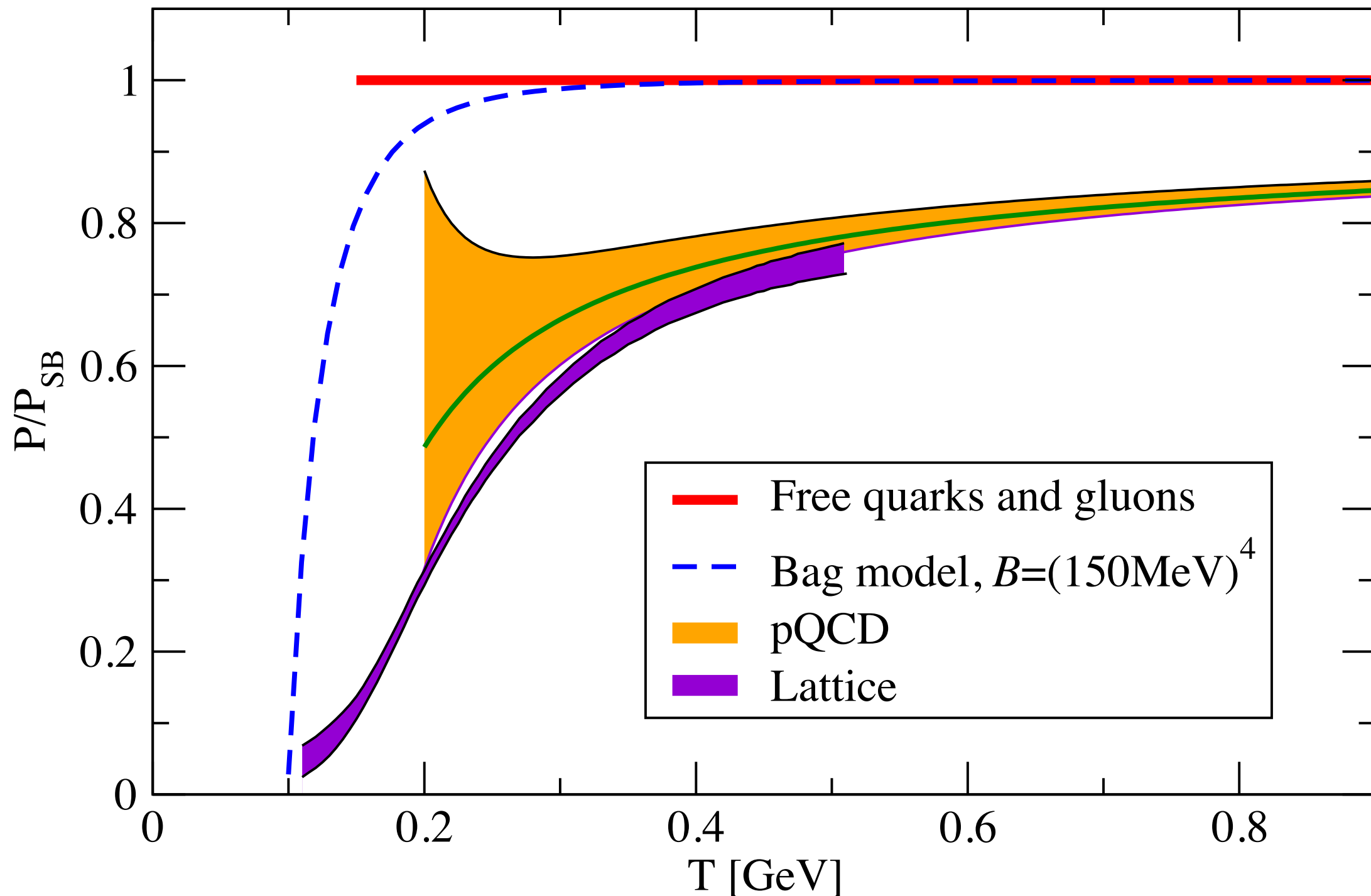
$$\bar{\Lambda} = \mu, 2\mu, 4\mu$$

Our result

with resummation



Analogy with high temperature case



Fraga, Kurkela, Vuorinen (2013)

Outline of the talk

- Introduction
- Hard Dense Loop (HDL) resummation and the setup of calculations
- The result for the EoS and its phenomenological implication
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Hard Thermal Loops (HTL)

- The problem of the gauge-dependent **gluon damping rate** at finite temperature (two scales: g & T) :

$$\gamma_g = a \frac{g^2 T}{8\pi}$$

Heinz, Kajantie, Toimela, ... (1987)

$a = 1$ (in Coulomb gauge), $a = -5$ (in covariant gauge)

- **Hard thermal loop (HTL) resummation:**

resum certain kinds of diagrams called HTLs, and use effective resummed propagator

Braaten, Pisarski (1990)

$a \simeq 6.635$ (both in Coulomb and covariant gauge)

Need of resummation

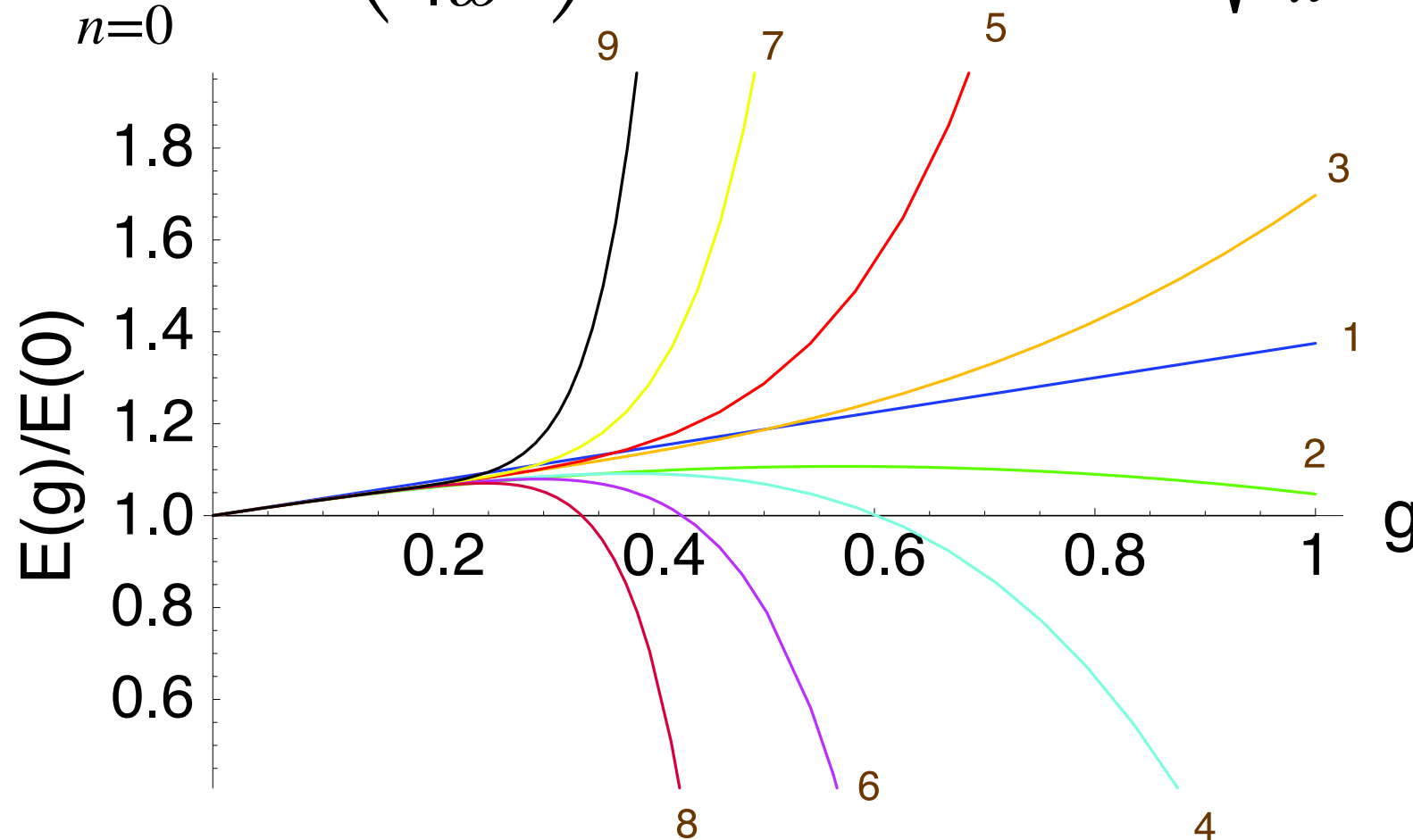
Prototypical example: an **anharmonic oscillator**

$$V(x) = \frac{\omega^2}{2}x^2 + \frac{g}{4}x^4$$

ground state energy:

Bender, Wu (1968)

$$E(g) = \omega \sum_{n=0}^{\infty} c_n^{\text{BW}} \left(\frac{g}{4\omega^3} \right)^n, \quad c_n^{\text{BW}} \xrightarrow{n \rightarrow \infty} (-1)^{n+1} \sqrt{\frac{6}{\pi^3}} 3^n \left(n - \frac{1}{2} \right) !$$



poor convergence...

Need of resummation

Variational perturbation theory (VPT):

Feynman, Kleinert, ... (1986)

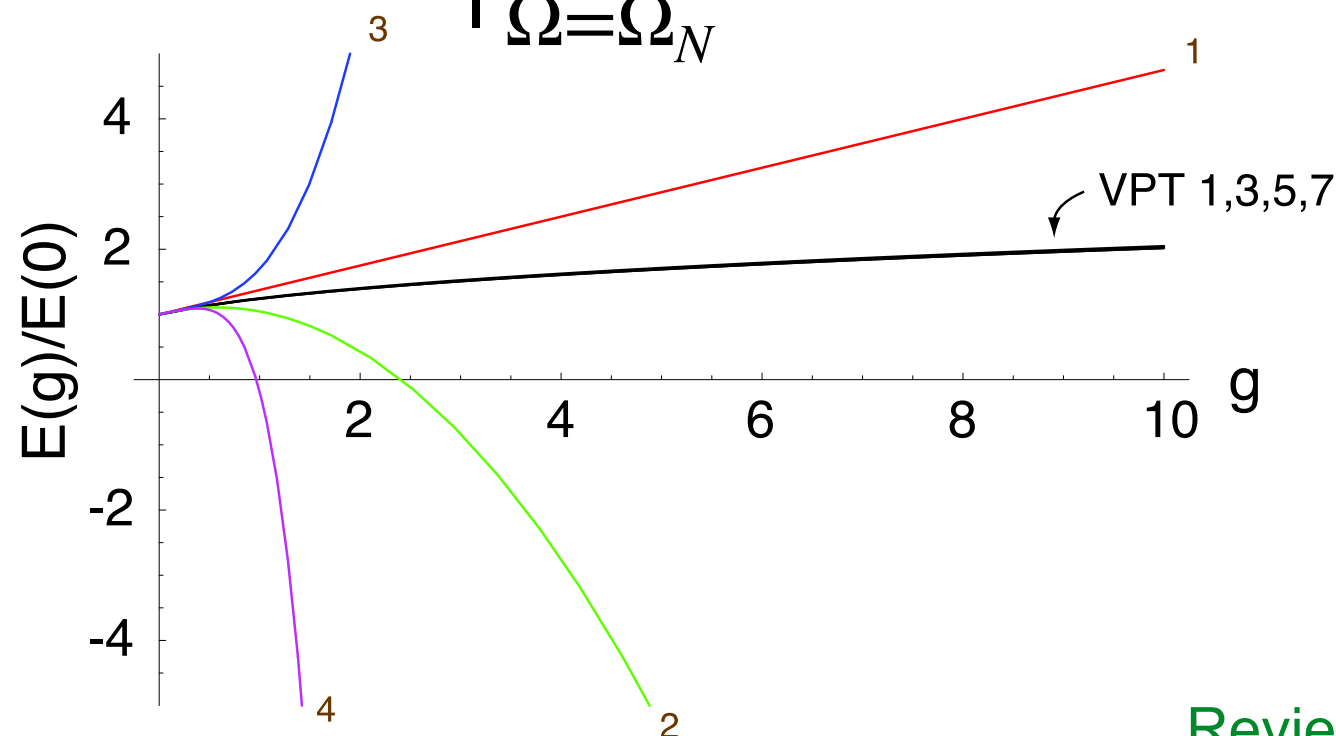
$$\omega^2 \rightarrow \Omega^2 + (\omega^2 - \Omega^2)$$

$$V(x) = \frac{\omega^2}{2}x^2 + \frac{g}{4}x^4 \rightarrow V(x) = \frac{\Omega^2}{2}x^2 + \frac{g}{4}(rx^2 + x^4)$$

$$(r := 2(\omega^2 - \Omega^2)/g)$$

minimize ground state energy w.r.t. Ω :

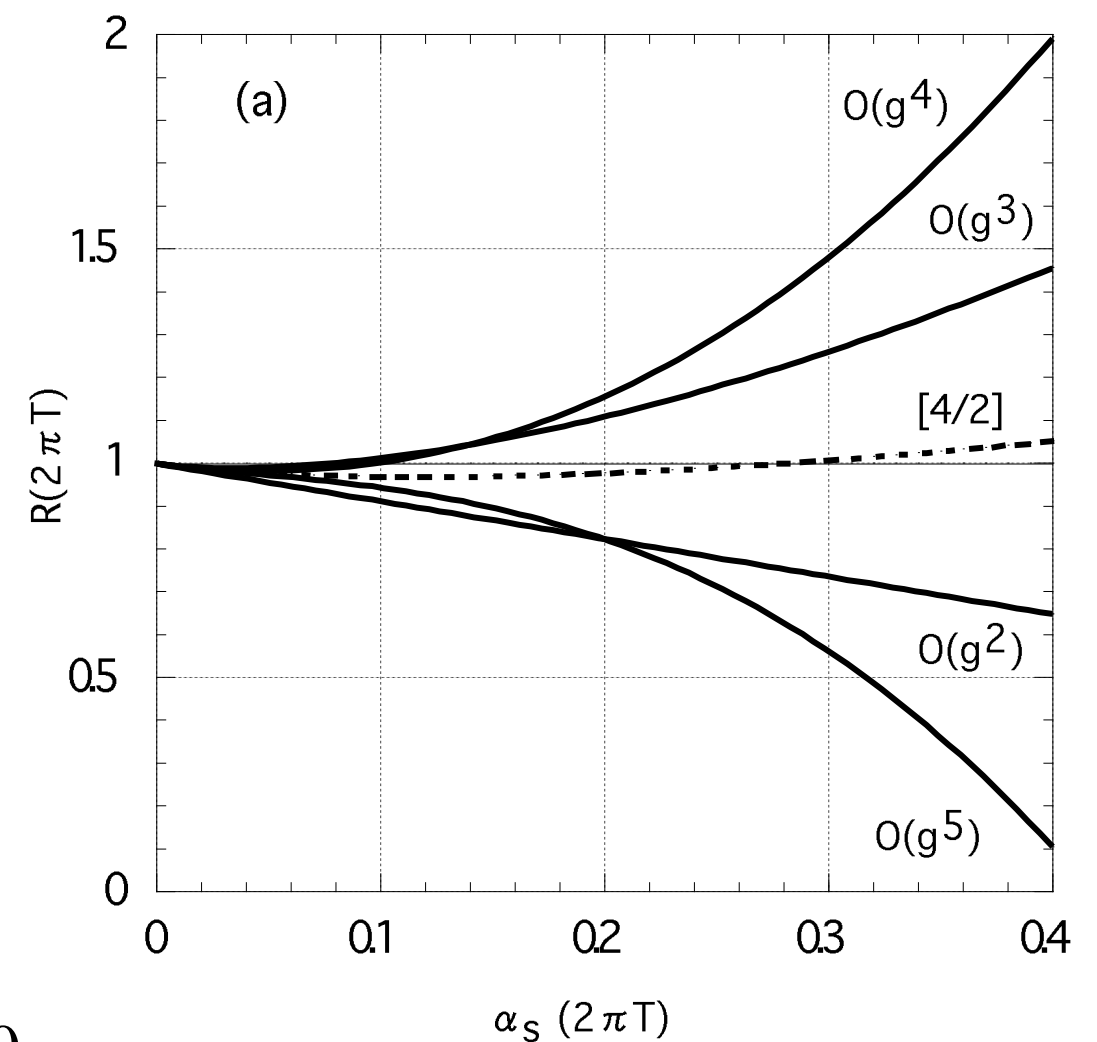
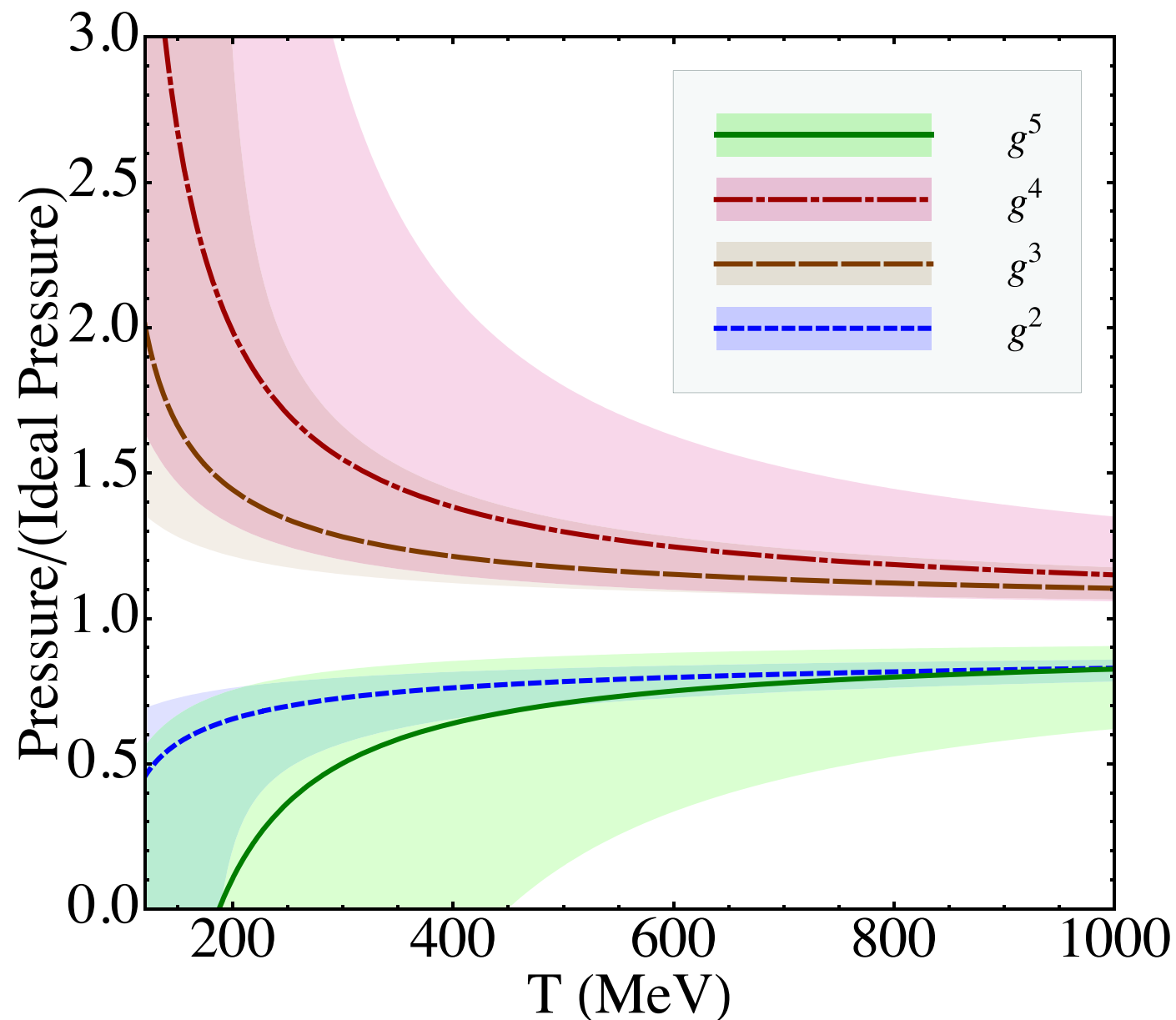
$$\left. \frac{\partial E_N}{\partial \Omega} \right|_{\Omega=\Omega_N} = 0, \quad (E_N(g, r) = \Omega \sum_{n=0}^N c_n(r) \left(\frac{g}{4\Omega^3} \right)^n)$$



Effective resummation
by VPT cures the
convergence problem

QCD thermodynamics at high T

- The same problem of poor convergence resides in the QCD EoS at high T :



Hatsuda (1997)

Review by Su(2012)

HTL perturbation theory

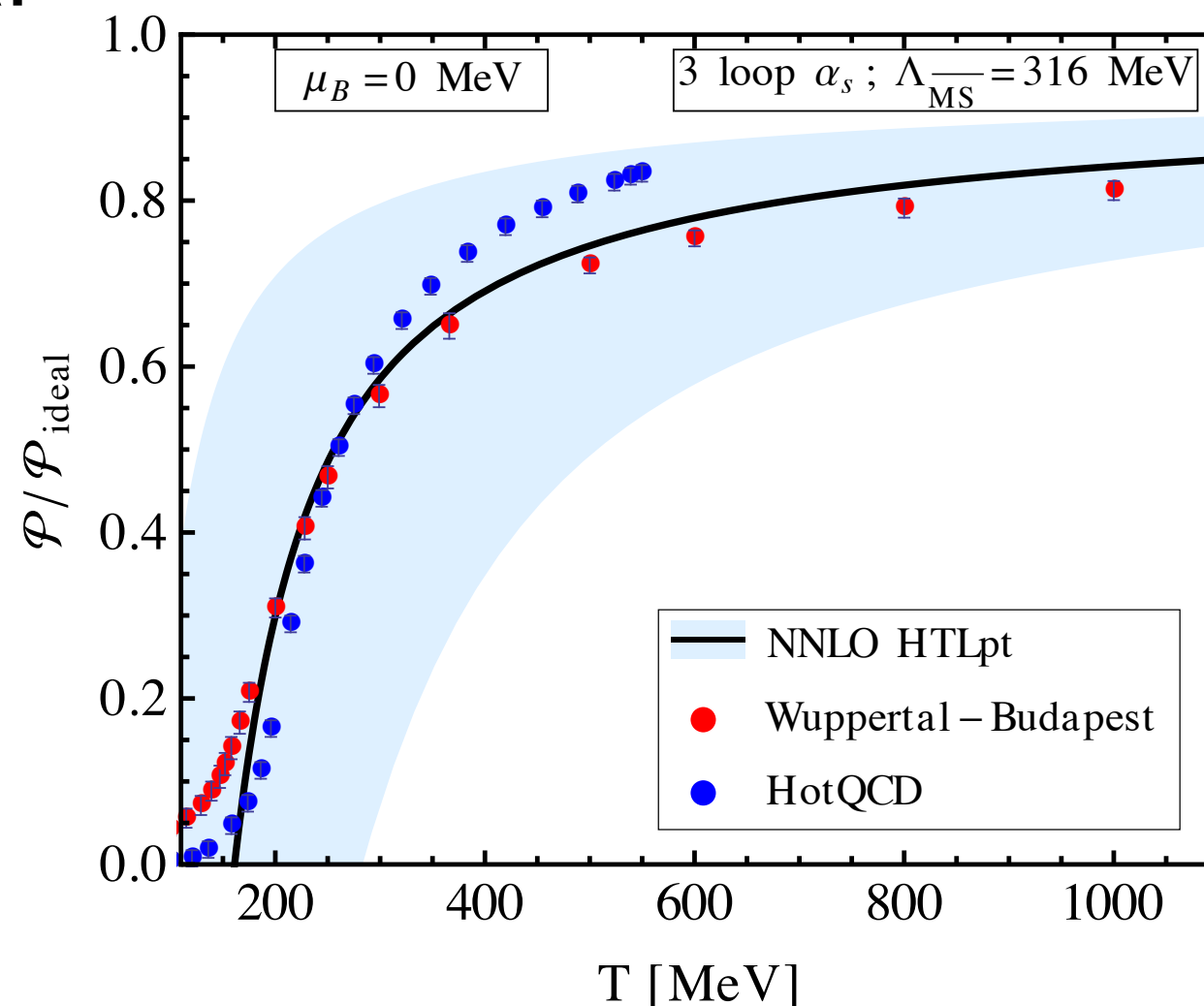
- HTL perturbation theory (sort of VPT) is one way to implement HTL resummation

Andersen, Braaten, Strickland, ... (1998)

cf) Pade: Hatsuda (1997), Optimized PT: Chiku, Hatsuda (1998),
EQCD: Blaizot, Iancu, Rebhan (2003), Laine, Schroder (2006)

- Hot QCD EoS with improved convergence confronting lattice data:

Haque, Bandyopadhyay, Andersen, Mustafa, Strickland, Su (2014)



Hard Dense Loops (HDLs)

- **Hard dense loops (HDLs):**

$T = 0$ and $\mu > 0$ counterpart of the HTLs

cf. thermal quark mass:

$$m_q^2 = \frac{g^2}{6} \left(\underline{T^2} + \frac{\mu^2}{\underline{\pi^2}} \right)$$

- Parallelism between $T \leftrightarrow \mu$ has been studied Manuel (1996)

- **EoS calculations at $T = 0$:**

- Baier, Redlich: hep-ph/9908372
- Andersen, Strickland: hep-ph/0206196

What we calculate here

Quark contribution to the pressure p :

$$p(\mu) = \text{Tr} \log S^{-1}$$

— HDL resummed full propagator

$$= \sum_{\{K\}} \ln \det [\not{k} - M_f - \Sigma(i\tilde{\omega}_n + \mu_f, k)]$$

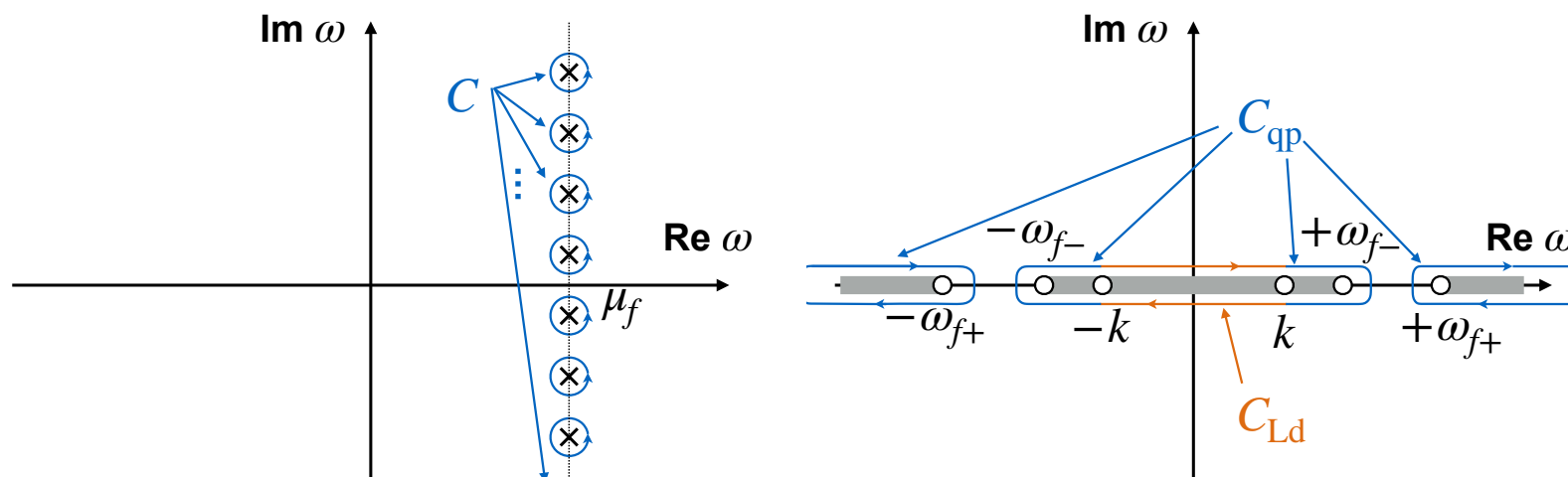
— Strange quark mass is included

Self-energy Σ in HDL approximation:

$$\Sigma \equiv \frac{m_q^2}{k} \gamma^0 Q_0 \left(\frac{k_0}{k} \right) + \frac{m_q^2}{k} \gamma \cdot \hat{k} \left[1 - \frac{k_0}{k} Q_0 \left(\frac{k_0}{k} \right) \right]$$

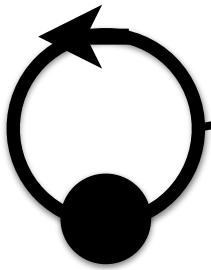
(Legendre function:
 $Q_0(x) = \frac{1}{2} \log \frac{x+1}{x-1}$)

Integration contour deformation:



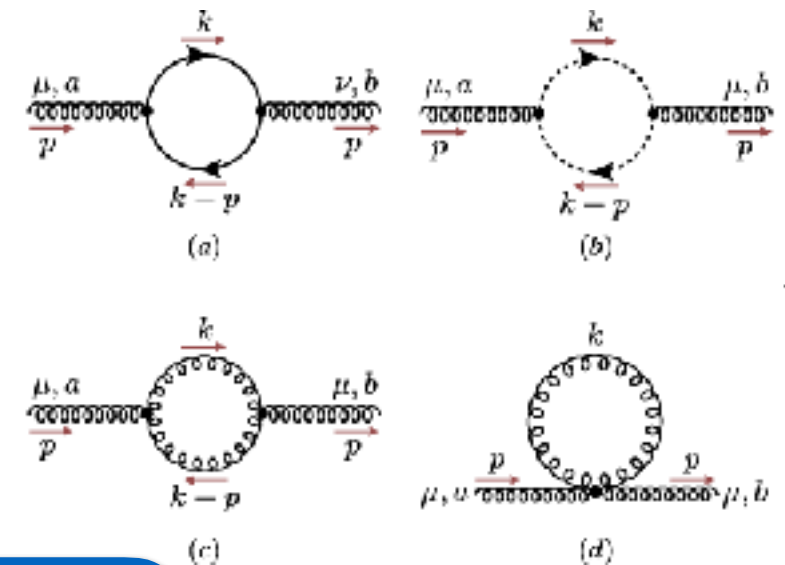
tedious calculation...

Diagrammatic argument

What we calculate:  full propagator

Hard dense loops:

higher order diagrams already
appears at the leading order

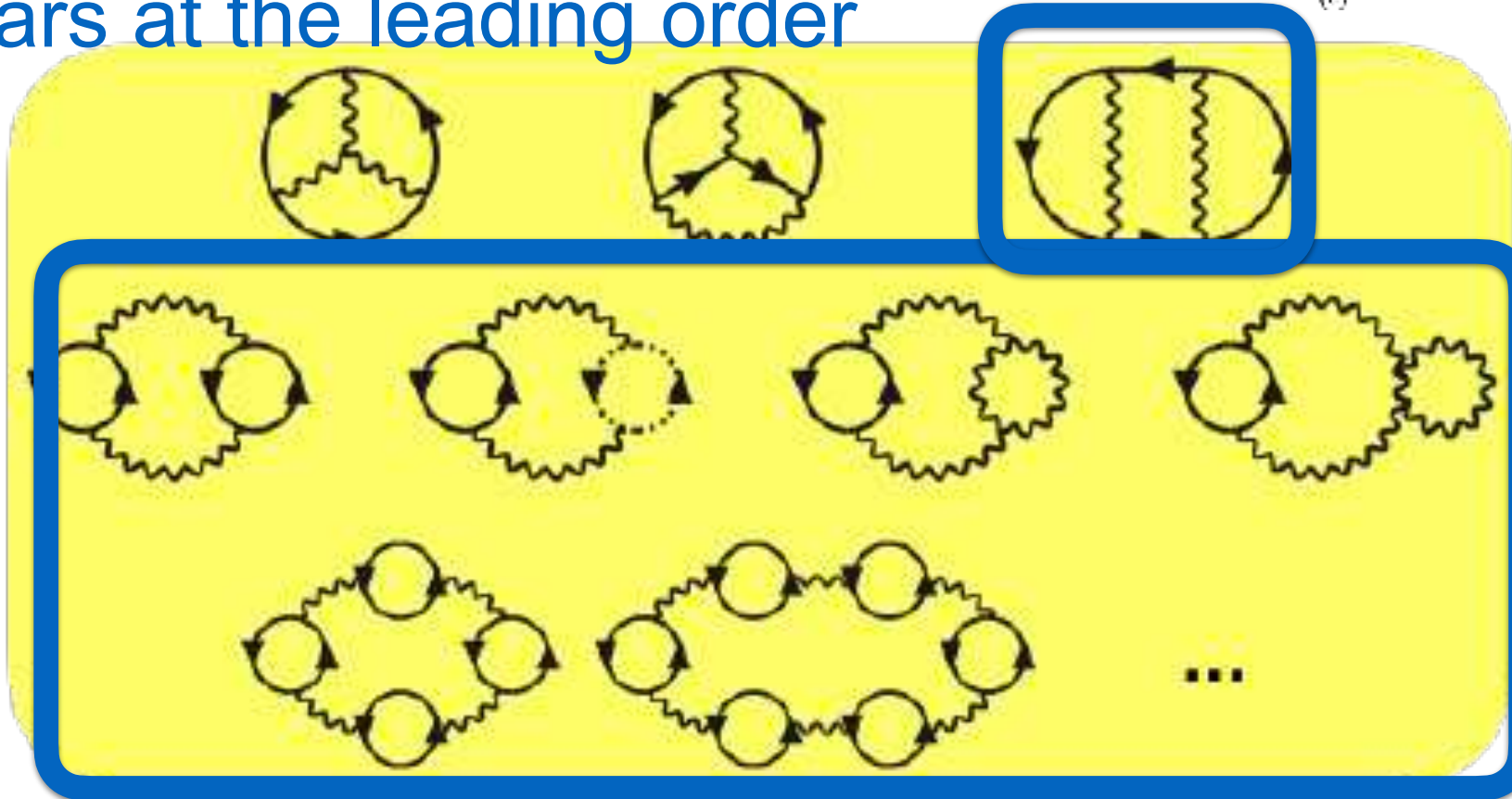


(5.23)

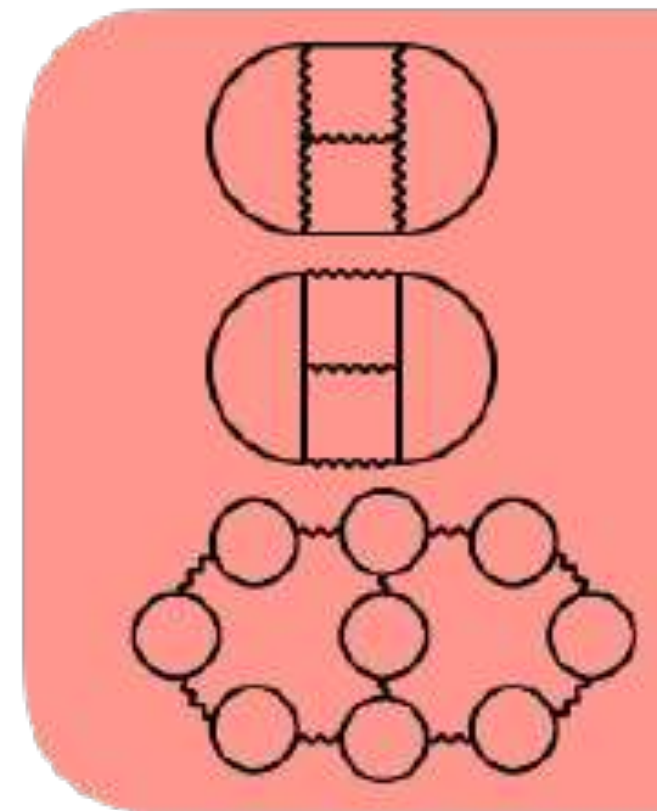
NNNLO



NLO



NNLO



In terms of the 2PI language

- **2PI (CJT) formalism:**

Cornwall,Jackiw,Tomboulis(1974)

Blaizot,Iancu,Rebhan (1999-)

$$p[D, S] = -\frac{1}{2}\text{Tr} \log D^{-1} + \frac{1}{2}\text{Tr} \Pi D \\ + \text{Tr} \log S^{-1} - \text{Tr} \Sigma S - \Phi_2[D, S]$$

(D : gluon full propagator)

$$\Phi[D, S] = -1/12 \text{ (triangle loop)} -1/8 \text{ (box loop)} +1/2 \text{ (self-energy)} + \dots$$

- **2PI formalism** employs propagator with the self-energy insertion $\rightarrow$ quasi-particle approximation
- May not be a systematic expansion in the coupling α_s

$\bar{\Lambda}$ (renormalization scale) dependence

- Use perturbative expression for screening mass m_q in Σ :

$$m_q^2 = \frac{4\alpha_s(\bar{\Lambda})}{3\pi} \mu^2 \quad \left(\Sigma \equiv \frac{m_q^2}{k} \gamma^0 Q_0 \left(\frac{k_0}{k} \right) + \frac{m_q^2}{k} \gamma \cdot \hat{k} \left[1 - \frac{k_0}{k} Q_0 \left(\frac{k_0}{k} \right) \right] \right)$$

- Renormalization scale $\bar{\Lambda}$ dependence enters the theory through the running coupling $\alpha_s(\bar{\Lambda})$:

$$\alpha_s(\bar{\Lambda}) = \frac{4\pi}{\beta_0 \log(\bar{\Lambda}^2 / \Lambda_{\overline{\text{MS}}}^2)} \times \left[1 - \frac{2\beta_1}{\beta_0^2} \frac{\log^2(\bar{\Lambda}^2 / \Lambda_{\overline{\text{MS}}}^2)}{\log(\bar{\Lambda}^2 / \Lambda_{\overline{\text{MS}}}^2)} \right]$$

- $\bar{\Lambda}$: **only undetermined const.**
set to the typical scale of system:

$$\bar{\Lambda} = 2\mu$$

$$(\Lambda_{\overline{\text{MS}}} = 378 \text{ MeV})$$

“uncertainty” of $\bar{\Lambda}$ is evaluated as:

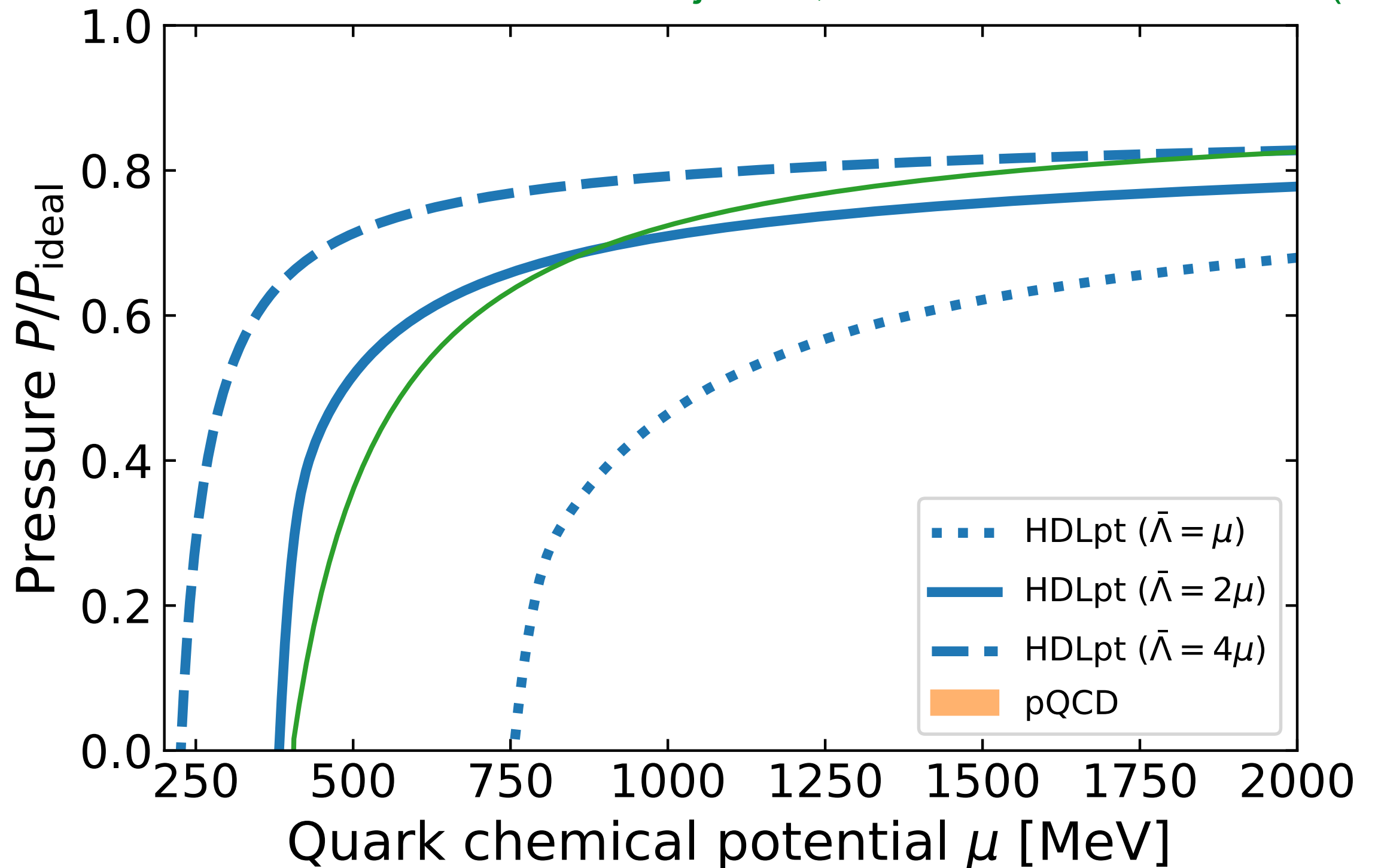
$$\bar{\Lambda} \in [\mu, 4\mu]$$

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Result from the HDL resummed QCD

Fujimoto, Fukushima: 2011.10891 (2020)

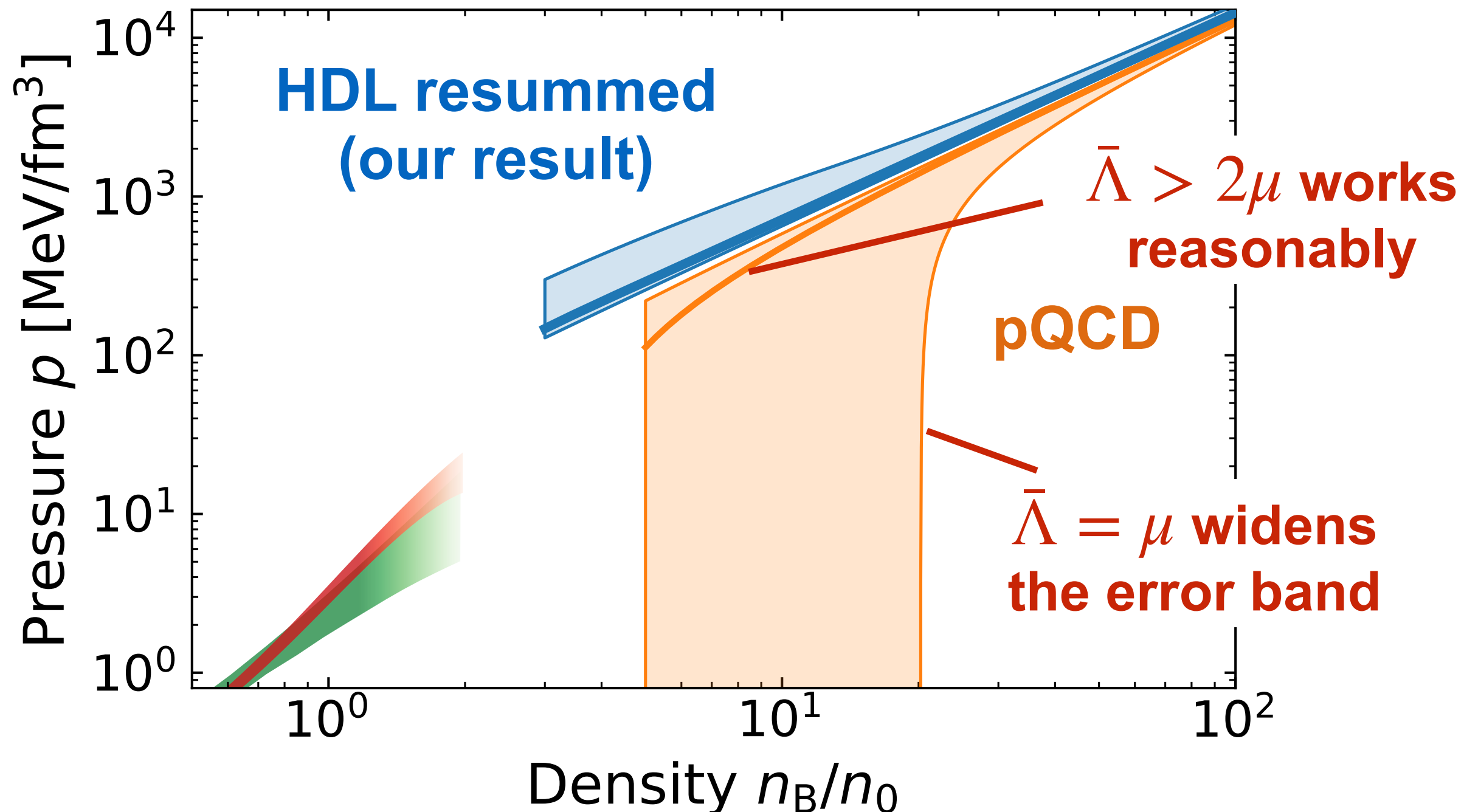


No convergence problem from the beginning.
Uncertainty seems not much improving...

Result from the HDL resummed QCD

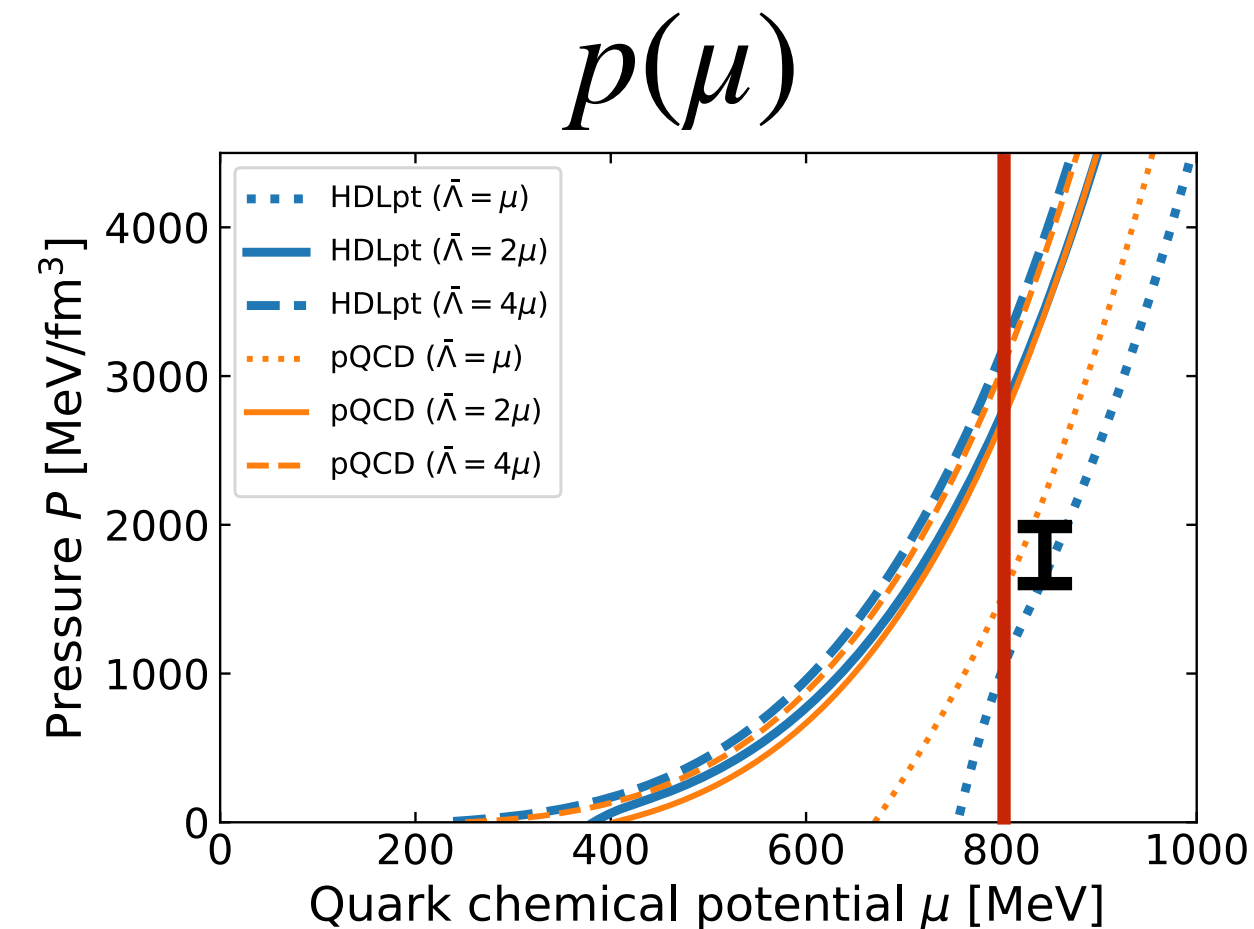
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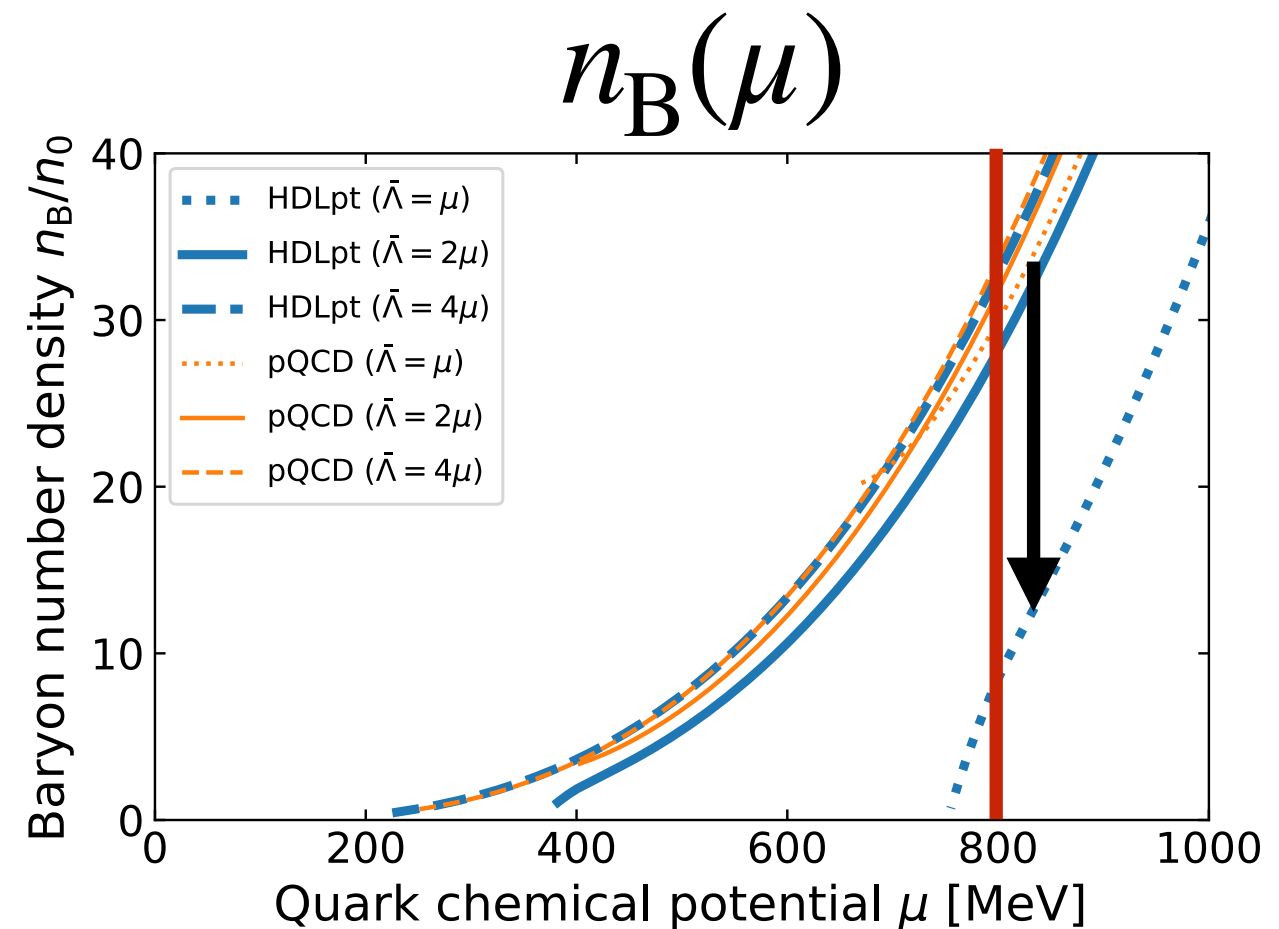
Heuristic argument

Fujimoto, Fukushima: 2011.10891 (2020)



Pressure does not differ
at constant μ

→ in HDL resummation,
the same value of p realizes at lower n_B
especially for $\bar{\Lambda} = \mu$

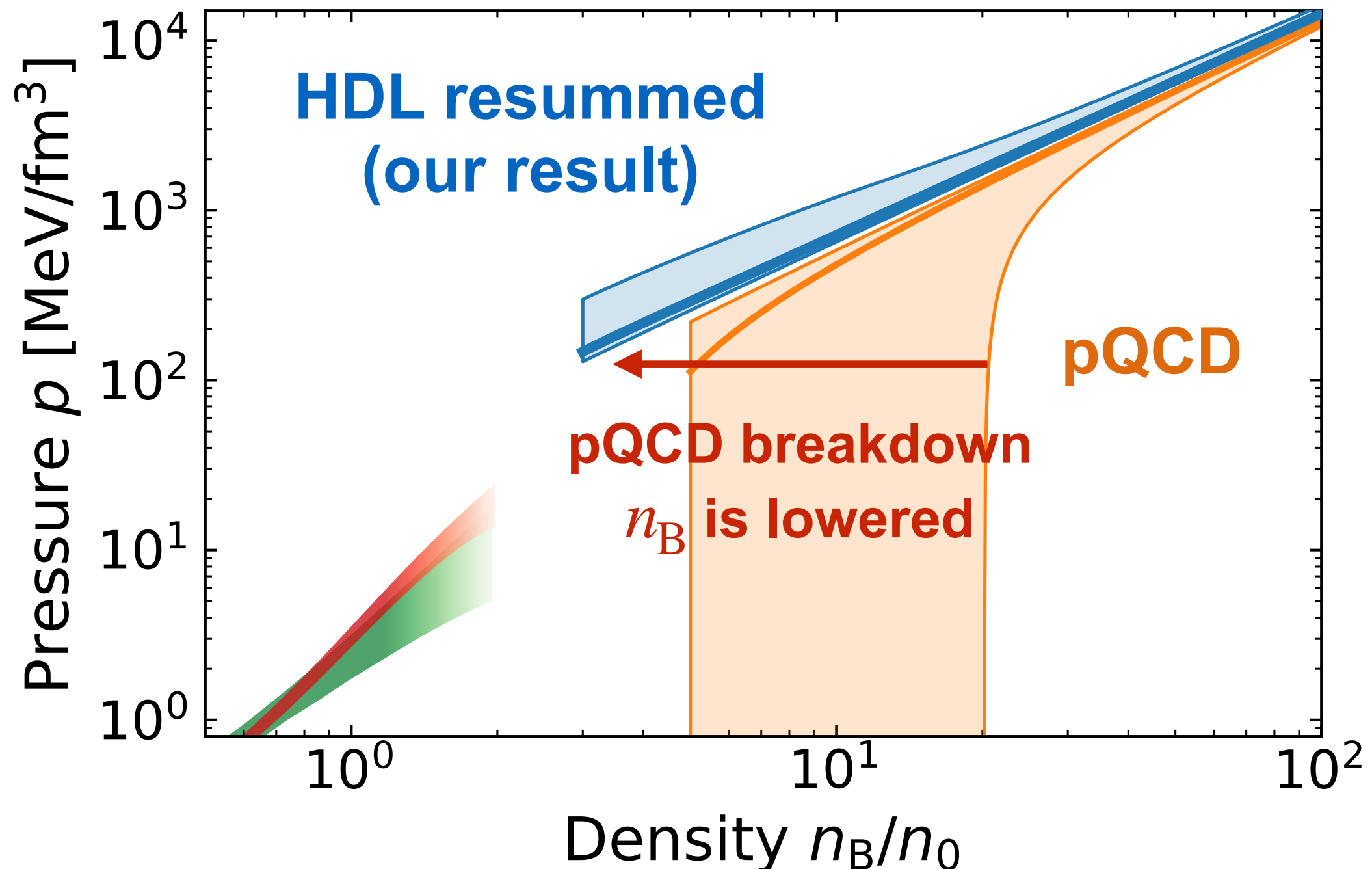


Density is screened in HDL
resummation at constant μ

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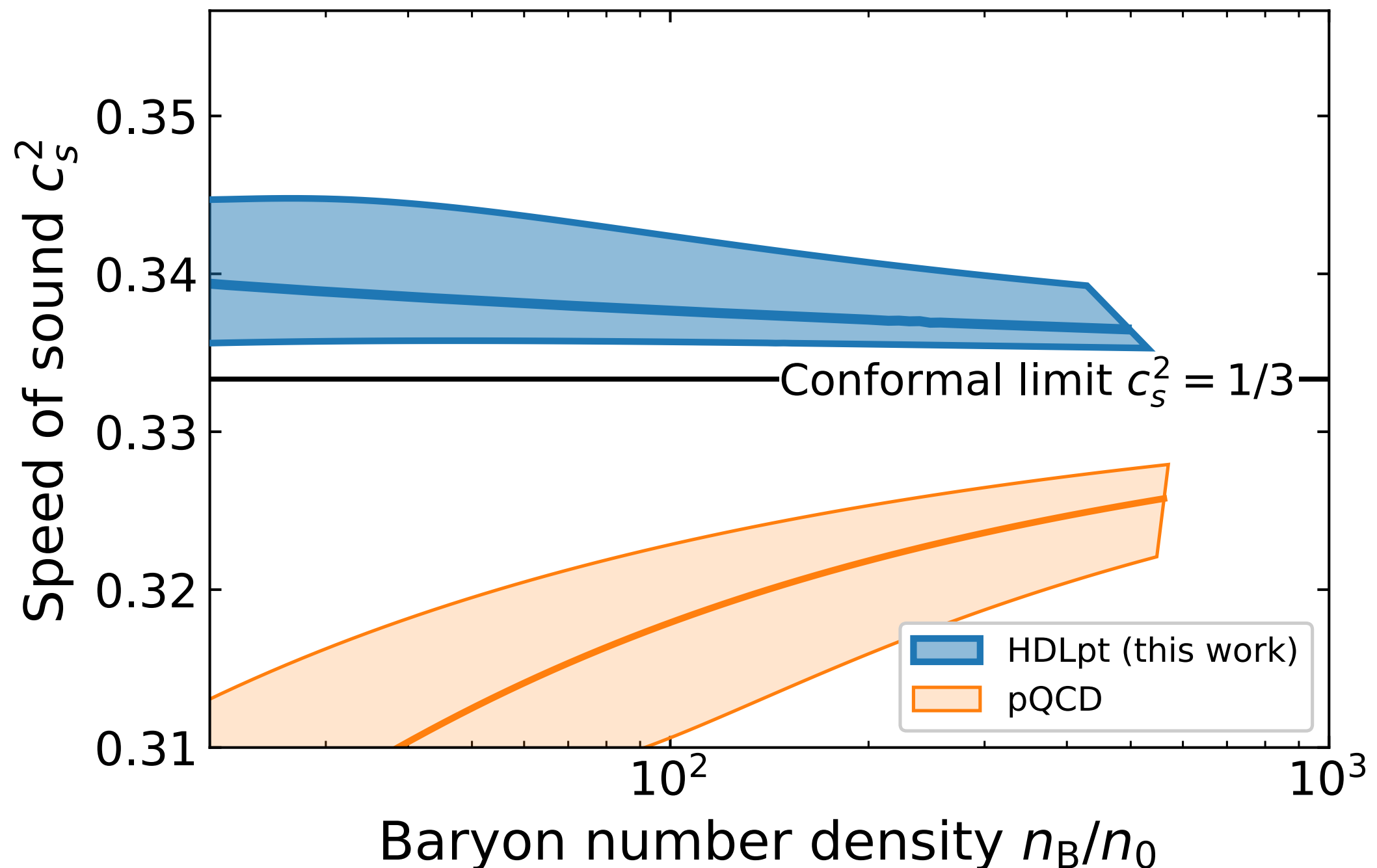
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Speed of sound

Fujimoto, Fukushima: 2011.10891 (2020)

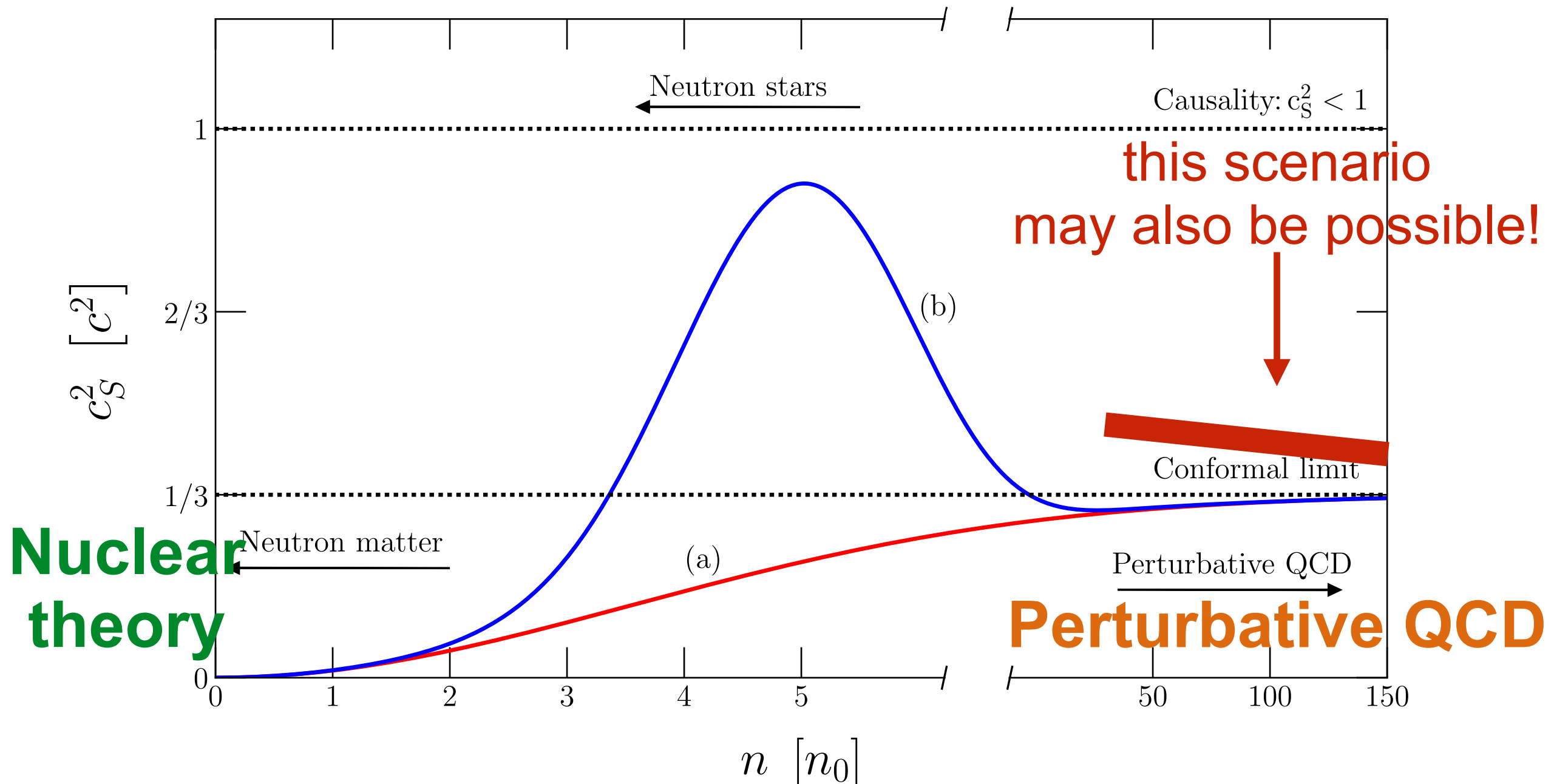
- Speed of sound: $c_s^2 = \partial p / \partial \varepsilon$;
asymptotic freedom: $c_s^2 \rightarrow 1/3$ (“**conformal limit**”)



Speed of sound

Bedaque, Steiner (2015)
Tews, Carlson, Gandolfi, Reddy (2018)

- The commonly shown graph for speed of sound

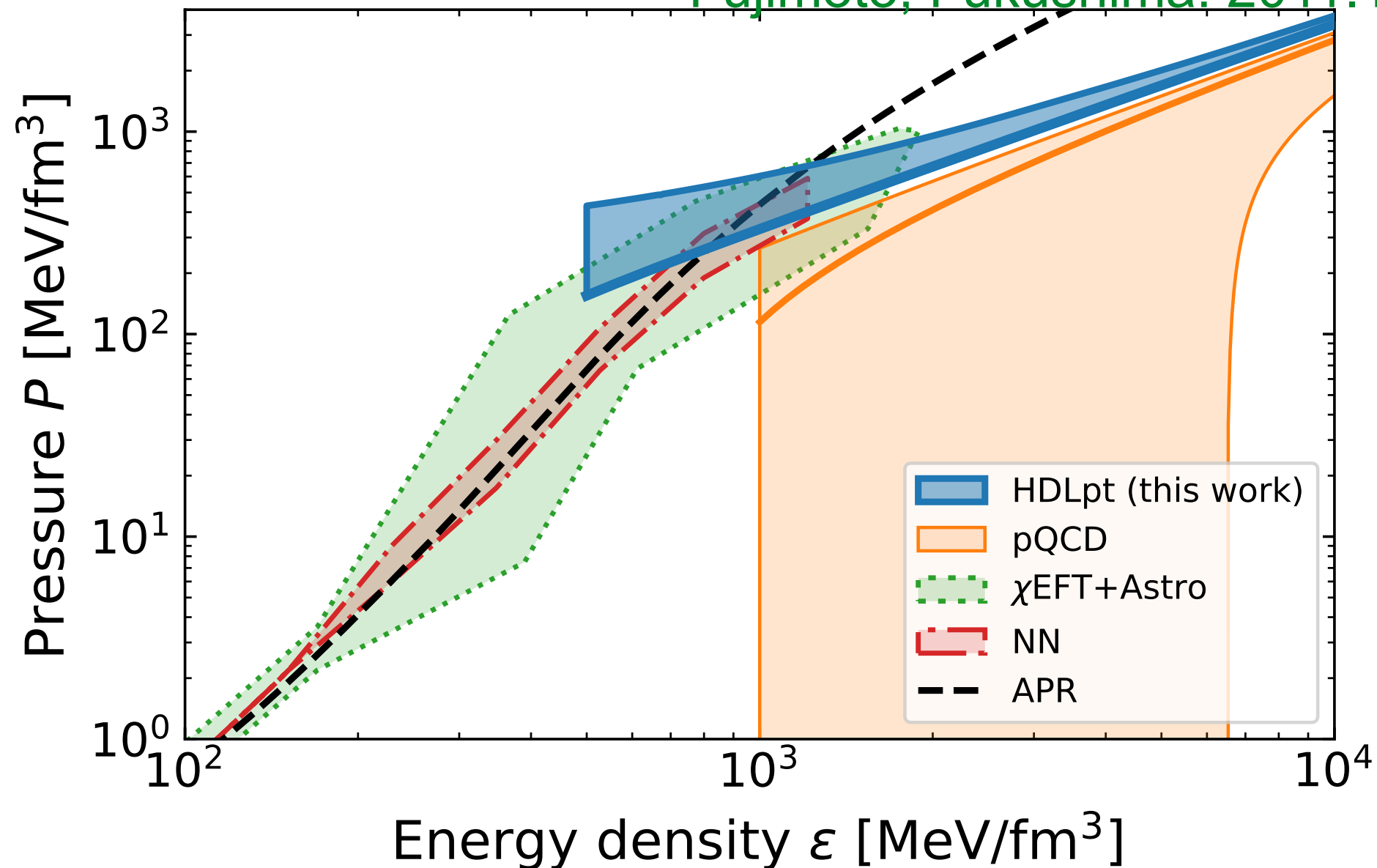


If $c_s^2 > 1/3$ at high density, then there must be a peak

Many discussions, e.g., Hippert, Fraga, Noronha (2021)

Smooth matching to the nuclear EoS?

Fujimoto, Fukushima: 2011.10891 (2020)



NN EoS extracted from X-ray NS data using neural network

Fujimoto, Fukushima, Murase (2017-)

χ EFT+Astro $n_B \leq n_0$: ChEFT EoS

$n_B > n_0$: extrapolation w/ the $2M_\odot$ pulsar constraint

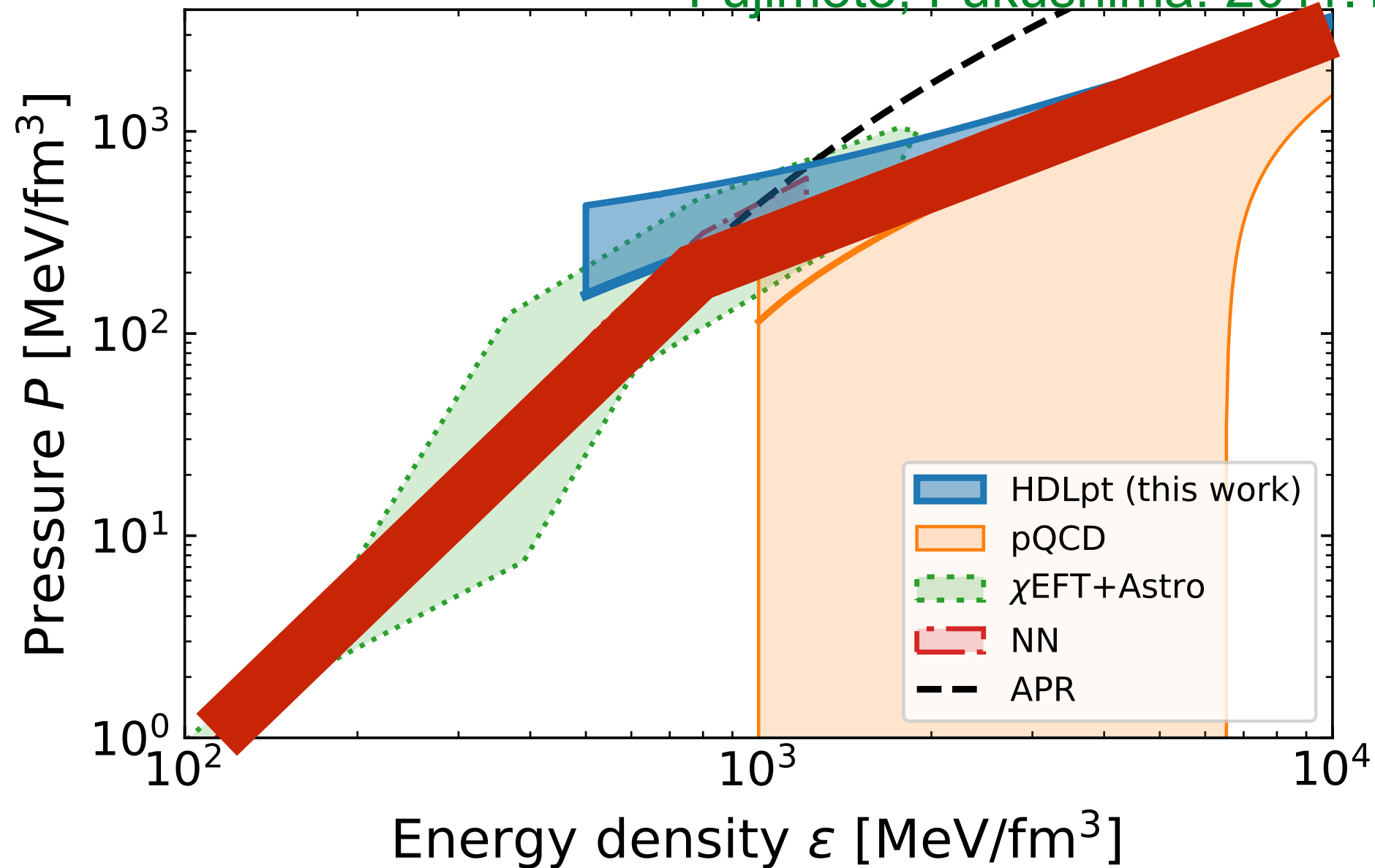
Hebeler, Lattimer, Pethick, Schwenk (2013)

APR conventional nuclear EoS

Akmal, Pandharipande, Ravenhall (1998)

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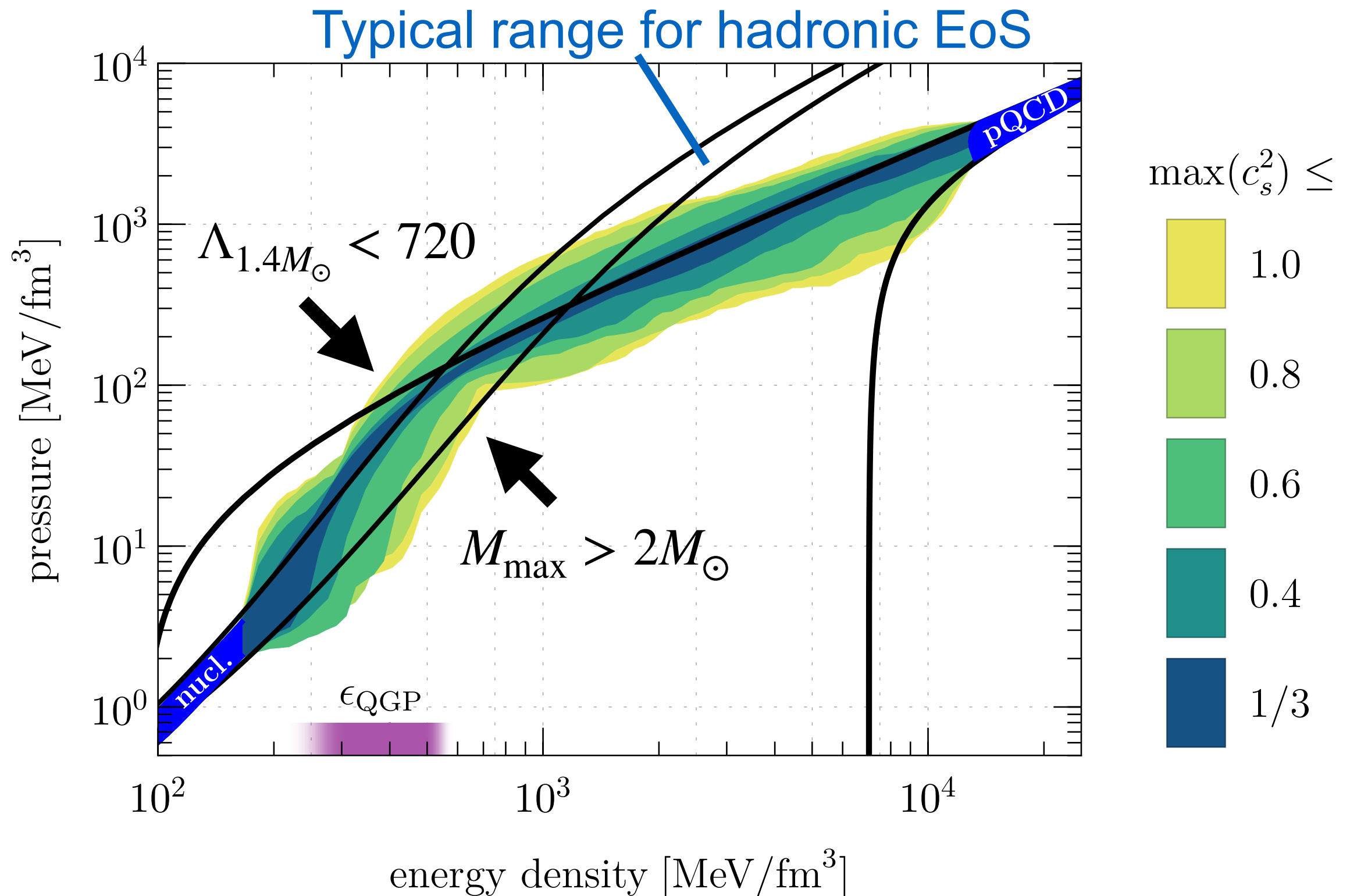
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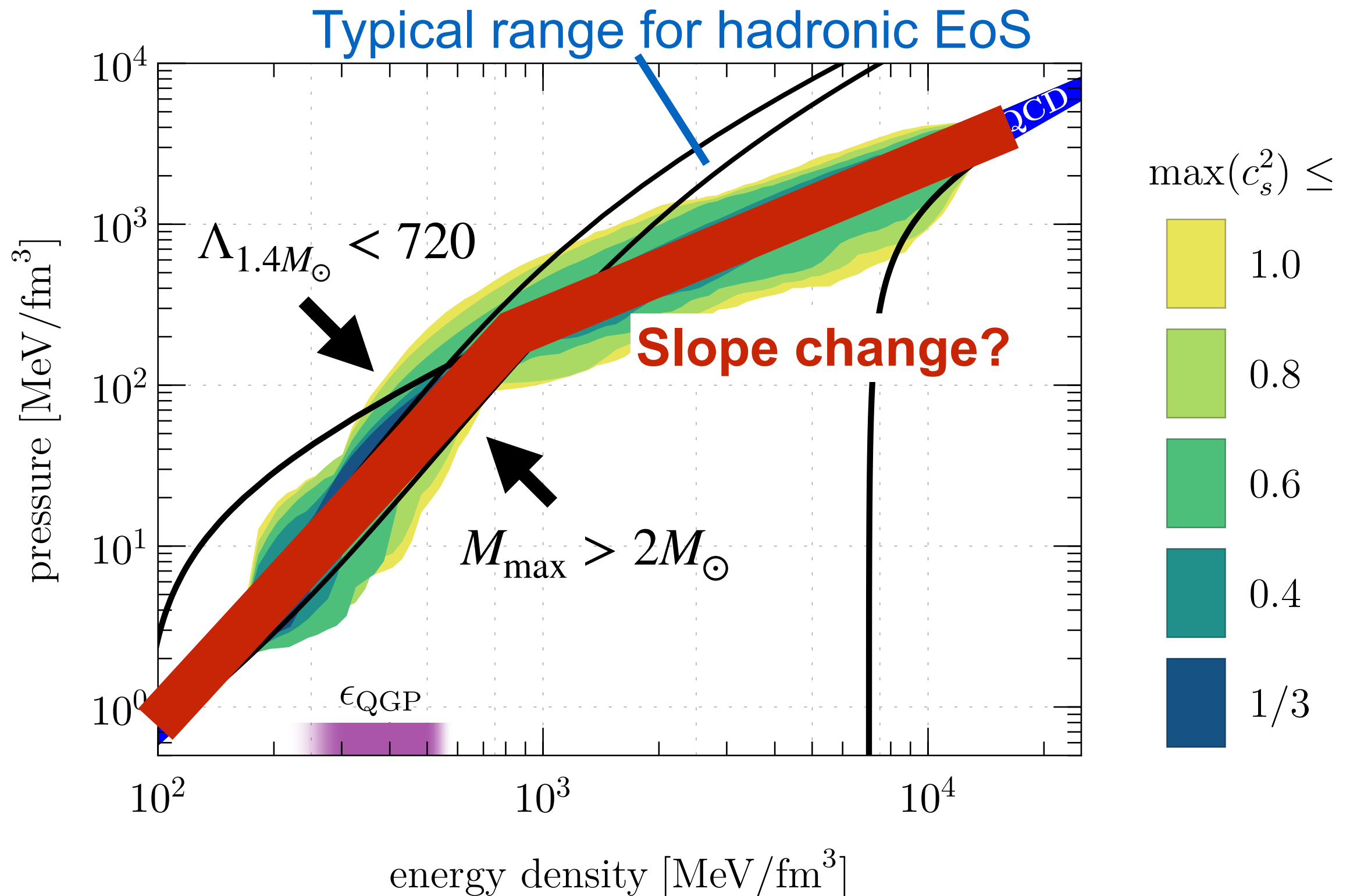
Observational constraint

Annala, Gorda, Kurkela, Nattila, Vuorinen (2019)



Observational constraint

Annala, Gorda, Kurkela, Nattila, Vuorinen (2019)



Discussion: conventional pQCD vs HDLpt

- mismatch already at the order of $\mathcal{O}(\alpha_s)$:

Baier, Redlich (1999)

$$\frac{p_{\text{HDLpt}}}{p_{\text{ideal}}} \simeq 1 - 2N_c \frac{m_q^2}{\mu^2} + \mathcal{O}\left(\frac{m_q^4}{\mu^4}\right) = 1 \boxed{-4\frac{\alpha_s}{\pi}} + \mathcal{O}(\alpha_s^2)$$
$$\frac{p_{\text{pQCD}}}{p_{\text{ideal}}} = 1 \boxed{-2\frac{\alpha_s}{\pi}} - \mathcal{O}(\alpha_s^2)$$

Resolving the $\mathcal{O}(\alpha_s)$ discrepancy

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Baier, Redlich (1999)

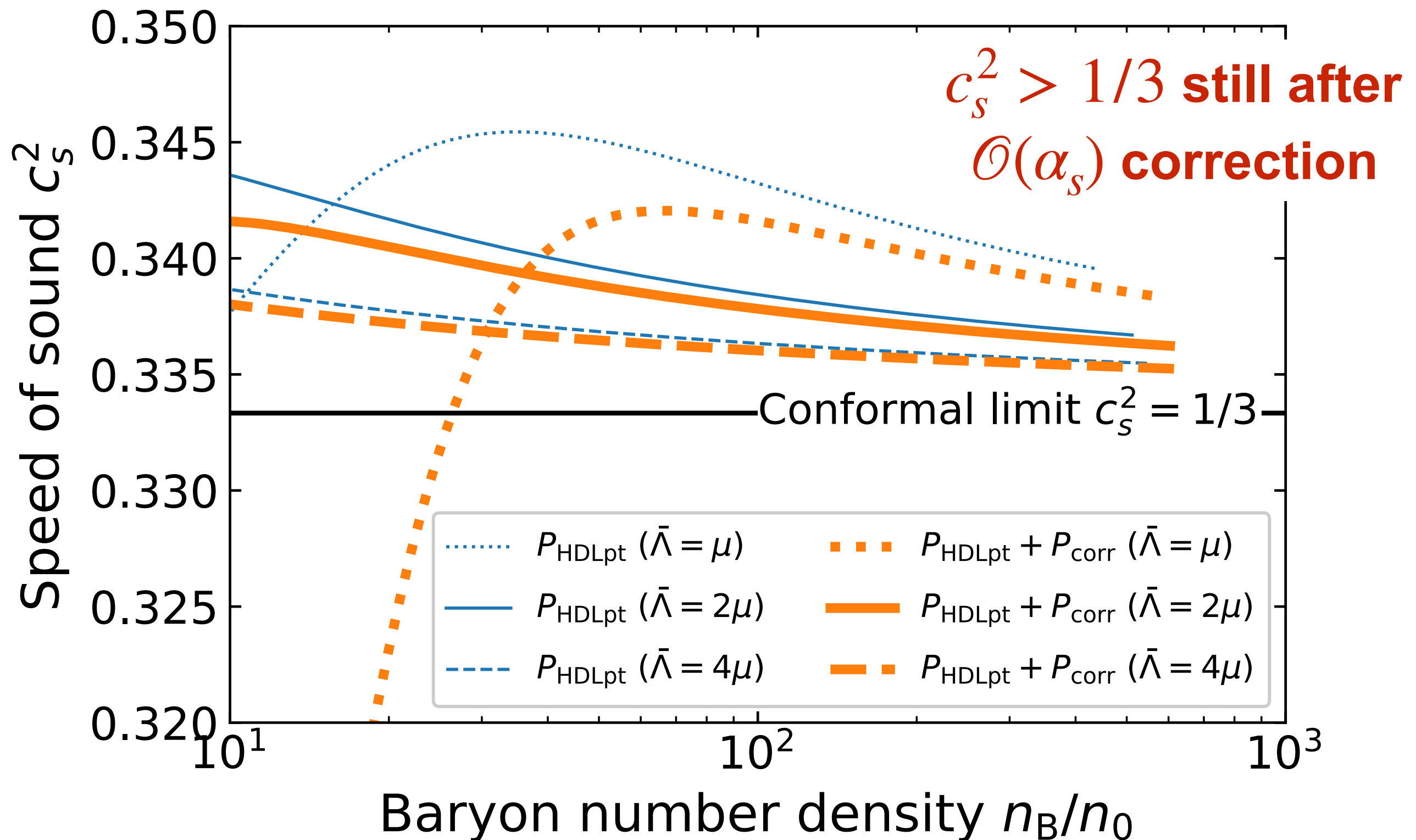
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$$\frac{p_{\text{pQCD}}}{p_{\text{ideal}}} = 1 \boxed{-2\frac{\alpha_s}{\pi}} - \mathcal{O}(\alpha_s^2)$$

- Add $p_{\text{corr}} = 2\frac{\alpha_s}{\pi} p_{\text{ideal}}$ to p_{HDLpt} to remedy the $\mathcal{O}(\alpha_s)$ discrepancy
- $c_s^2 < 1/3$ in the pQCD calculation is attributed to negative β -function at $\mathcal{O}(\alpha_s)$, so we care about $\mathcal{O}(\alpha_s)$ discrepancy

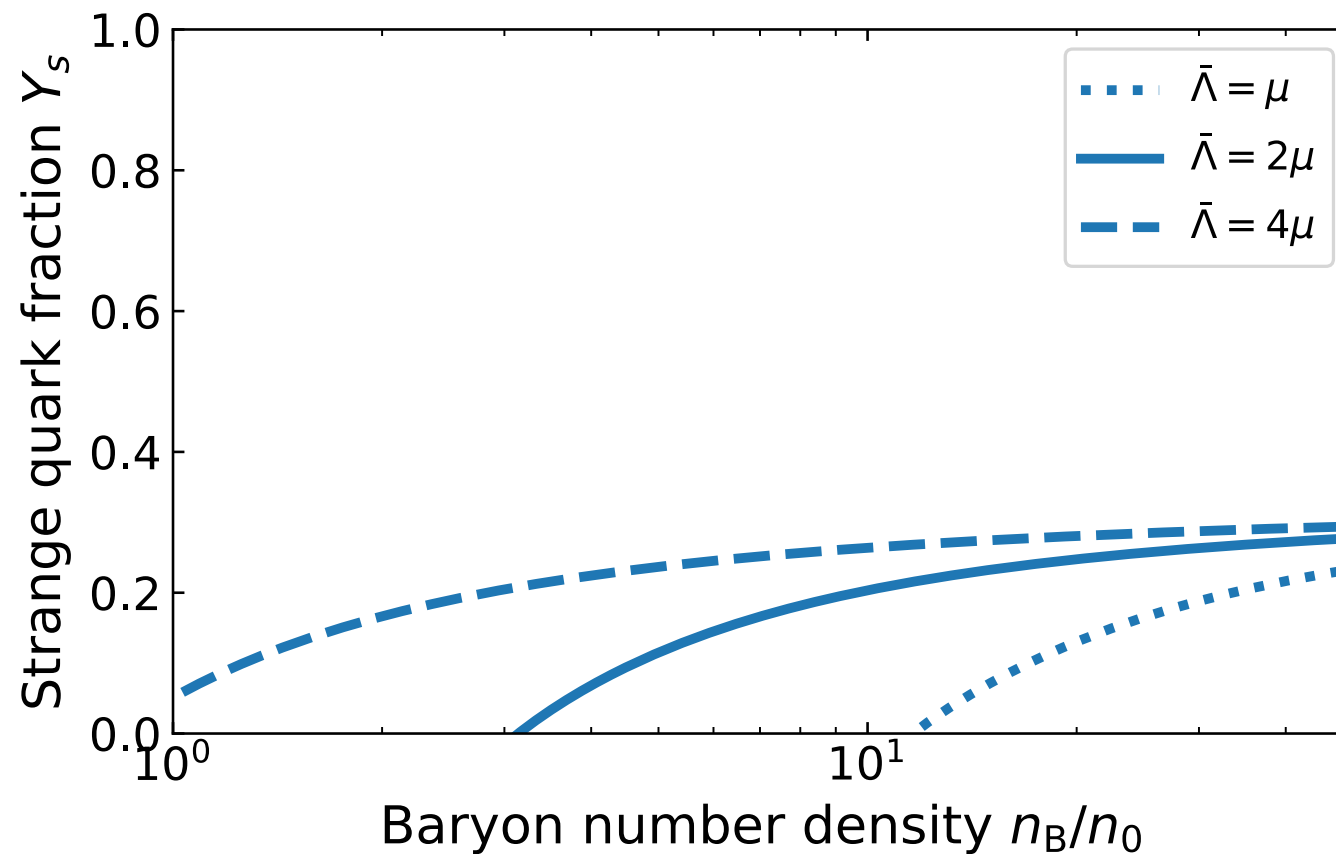
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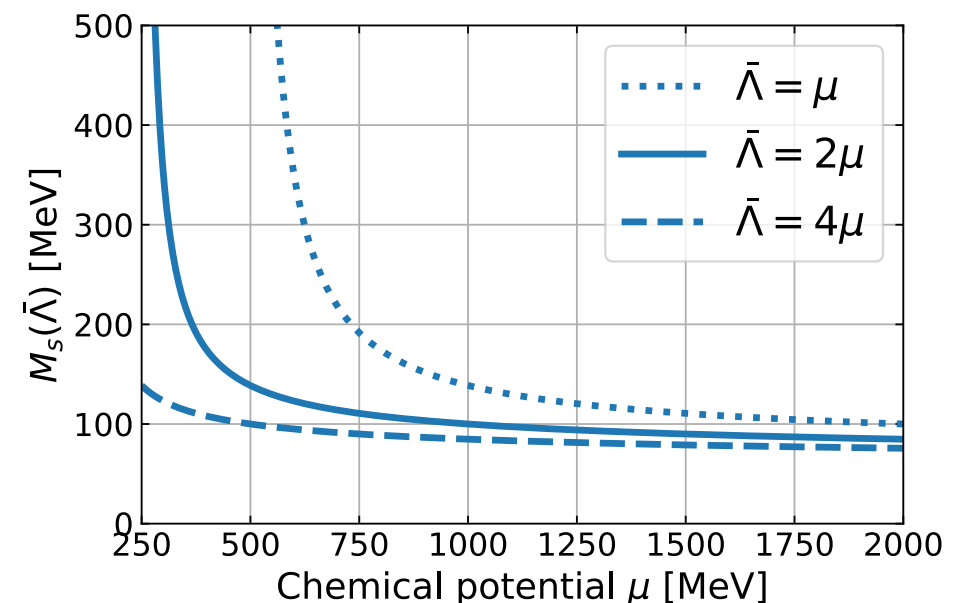
Discussion: strangeness content

- Fraction of the strange quarks:



large scale
dependence...

- Threshold strongly depends on the running of the current mass of strange quarks:



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Summary and outlook

- Resumming hard dense loops: systematic reorganization of perturbation theory, convergence improved
- The result turned out to extend the pQCD applicability down to the realistic density in neutron stars
- Several issues to be explored:
 - * Deeper reason why uncertainty is smaller?
 - * Evaluating higher order 2PI skeleton diagrams?
 - * Multi-pronged *ab initio* approach to the EoS:
QCD + ChEFT + NS observation + (QHC?) + ...