iTHEMS Theoretical Physics Seminar @ RIKEN 2021/8/16

Application of AdS/CFT to nonequilibrium phenomena in external electric fields Shunichiro Kinoshita (Chuo University)

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- AdS/CFT correspondence
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JHEP 09 (2014) 126 [arXiv:1407.0798] Phys.Lett.B 746 (2015) 311-314 [arXiv:1408.6293] Nucl.Phys.B 896 (2015) 738-762 [arXiv:1412.4964] K.Hashimoto, SK, K.Murata, T.Oka

Application 2: rotating electric field

JHEP 05 (2017) 127 [arXiv:1611.03702] K.Hashimoto, SK, K.Murata, T.Oka JHEP 06 (2018) 096 [arXiv:1712.06786] SK, K.Murata, T.Oka

### ADS/CFT CORRESPONDENCE

# AdS/CFT correspondence

- A duality relating a classical gravity in (n+1)-dim. anti-de Sitter (AdS) space and a strongly correlated conformal field theory (CFT) in n-dim.
  - Holography, the gauge/gravity duality



- The classical dynamics of gravity corresponds to the quantum physics of strongly correlated gauge theory
  - General relativity could describe strongly corrected quantum systems (with finite temperature), which are too difficult to solve.
  - QCD, Quark-gluon plasma (QGP), condensed matter physics, …

## Anti-de Sitter(AdS) space

- Vacuum solution of the Einstein eq. with negative cosmological constant
- At infinity there is timelike boundary



Mathematical statement

E

$$Z_{\text{boundary}}[J] = \exp(iS_{\text{bulk}}[\Phi_{\text{cl}}])$$
$$J(x) = \Phi_{\text{cl}}(r, x)|_{r=\infty}$$

Partition function of CFT: 
$$Z_{\text{boundary}}[J] \equiv \int \mathcal{D}\phi \exp\left[iS_{\text{CFT}} + i\int d^4x J\mathcal{O}\right]$$
  
 $\langle \mathcal{O}(x) \rangle = -i \frac{\delta}{\delta J(x)} \log Z[J] = \frac{\delta S_{\text{bulk}}[\Phi_{\text{cl}}]}{\delta \Phi_{\text{cl}}|_{r=\infty}}$   
×. Gravitational field  $\langle T_{\mu\nu}^{\text{CFT}} \rangle = \frac{\delta S_{\text{bulk}}[g_{\mu\nu}^{\text{cl}}]}{\delta g_{\mu\nu}^{\text{cl}}|_{r=\infty}} \qquad g_{\mu\nu}(r,x) = \eta_{\mu\nu} + \frac{1}{r^4} \langle T_{\mu\nu}^{\text{CFT}} \rangle + \cdots$   
field theory gravity  $ds^2 = \frac{dr^2}{r^2} + r^2 g_{\mu\nu}(r,x) dx^{\mu} dx^{\nu}$ 

### Brief manual

• Field theory

J(x)

External source



Classical gravity

$$\Phi(r,x)|_{r=\infty} = J(x)$$

Boundary condition of bulk field

Solve classical EoM



**Bulk solution** 

 $\Phi(r,x)$ 

Expectation value of response

Expand it around AdS boundary ( $r = \infty$ )

$$\Phi(r,x) = J(x) + \frac{\langle O(x) \rangle_J}{r^{\Delta}} + \cdots$$

### D3/D7 BRANE SYSTEM

### String theoretical viewpoint



Strongly coupled gauge theory corresponds to classical gravity A new type brane will be added for quark degrees of freedom

# Holographic QCD constructed by D3/D7

Karch, Katz (2002), Grana, Polchinski (2002), Bertolini et al. (2002)



# Phase transition in the D3/D7 system

 If we introduce a worldvolume gauge field living on the D7 brane or a black hole in bulk spacetime, topology of the brane configuration can change

Electric field or black hole become larger

Critical embedding

Minkowski embedding





Black hole embedding

Fluctuations are normal modes with real frequencies (confined)

Fluctuations are quasi-normal modes with complex frequencies (absorbed into a horizon)

### Confinement/deconfinement in the meson sector Mateos, Myers, Thomson (2006, 2007)

• If one includes electric field or finite temperature in the system, the phase transition occurs

### Gravity side

Black hole in the bulk or Gauge field on the brane

The brane is bending

$$w(\rho) \sim m + \frac{c}{\rho^2} + \cdots,$$
  
 $a_x(\rho) \sim -E_x t + \frac{j_x}{2\rho^2} + \cdots,$ 

The brane intersects a horizon or not

The fluctuations dissipate or are confined (quasi-normal modes or normal modes)

Gauge theory side

Finite temperature in the gluon sector or Finite electric field

Expectation values change

 $\langle \bar{\psi}\psi 
angle \propto c\,$  : quark condensate

 $\langle ar{\psi} \gamma_\mu \psi 
angle \propto j_\mu$  : electric current

*m*:quark mass,  $E_{\chi}$ :electric field

Deconfinement or confinement

Mesons are unstable or stable

### Electric field case

Schwinger effect

Karch, O'Bannon (2007) Erdmenger, Meyer, Shock (2007) Albash, Filev, Johnson, Kundu (2007)



Beyond the critical electric field, an effective horizon emerges on the brane The electric current becomes non-zero value = Schwinger effect

### APPLICATION1: ELECTRIC FILED QUENCH

### Time-dependent phenomena

- We want to consider non-equilibrium phenomena beyond steady sates
  - Suddenly applied electric fields



### Our setup: time-dependent electric field

- Hashimoto, SK, Murata, Oka JHEP 09 (2014) 126
- Bulk spacetime

- AdS<sub>5</sub>× S<sup>5</sup>  
$$ds^{2} = \frac{1}{z^{2}} \left[ -dV^{2} - 2dVdz + dx^{2} + d\vec{x}_{2}^{2} \right] + d\phi^{2} + \cos^{2}\phi d\Omega_{3}^{2} + \sin^{2}\phi d\psi^{2}$$

- D7-brane The brane is symmetric in  $\vec{x}_3$ ,  $\Omega_3$ -directions
  - Embedding function :

$$V = V(u, v), \quad z = Z(u, v), \quad \phi = \Phi(u, v), \quad \psi = 0$$

- Gauge field :  $2\pi \alpha' A_a dy^a = a_x(u,v)dx$ 

Boundary conditions of  $a_x$  at the AdS boundary = electric field in the boundary theory



# Equations of motion of the brane



Coordinate conditions:  $C_1 \equiv h_{uu} + Z^2 (\partial_u a_x)^2 = 0$ , Double-null coordinates  $C_2 \equiv h_{vv} + Z^2 (\partial_v a_x)^2 = 0$ 

$$\partial_u \partial_v \mathbf{X} + \mathbf{f}(\mathbf{X}, \partial_u \mathbf{X}, \partial_v \mathbf{X}) = 0 \quad \mathbf{X} = (V, Z, \Phi, a_x)$$

2-dimensional non-linear wave equations on an effective metric

Effective metric:  $\gamma_{ab} \equiv h_{ab} + f_{ac} f_{bd} h^{cd}$ 

## Numerical Results

- Dynamical phase diagram
  - We found three cases depending on two model parameters:  $E_f$  and  $\Delta V$  electric field <sup>12</sup>



### Case 1 super-Schwinger-limit

• Strong electric field ( $E_f = 2.0, \Delta V = 0.50$ )



- The meson sector is deconfined by the Schwinger effect
- The system has been relaxed and thermalized
  - The effective horizon emerges on the worldvolume

### Thermalization time

· We can estimate thermalization time explicitly



### Case 2 sub-Schwinger-limit

• Weak electric field ( $E_f = 0.10, \Delta V = 2.0$ )



- Normal modes of fluctuations of the brane and the gauge field
   ⇔ discrete spectrum of meson
- "beat" ⇔ meson mixing
  - The Stark effect leads to splitting of degenerate mass spectrum

### Case 3

### Non-equilibrium deconfinement

- We found that sufficiently rapid quench causes the deconfinement transition even below the critical electric field
  - Notice that deconfinement is defined by divergence of redshift factor measured at the AdS boundary



The deconfinement time becomes discretely longer as the electric field is weaker

### "Turbulence" on the brane

### What is happening on the brane?

- The fluctuation caused by the quench at the boundary is amplified during coming and going on the brane
- After reflecting several times, a strongly red-shifted region emerges at the center and then a naked-singularity will form on the brane

### Worldvolume of D7 brane



It seems to be similar to "AdS turbulent instability" !

Discreteness of the deconfinement time = number of the reflections Deconfine (divergence of the redshift factor) = singularity formation

### Profile of the D7-brane in the bulk



### Spectral analysis

Hashimoto, SK, Murata, Oka arXiv:1408.6293, 1412.4964

- We decompose the brane fluctuations into the mode functions in linear perturbations
  - fluctuations of the brane and gauge field:



$$e_n(z) \equiv N_n z^2 F(n+3,-n,2;z^2)$$
  
Mode functions

$$S_{\rm D7} \propto \int d^4x \int_0^1 dz \frac{1-z^2}{2z} [(\partial_t \chi)^2 - m^2 (1-z^2) (\partial_z \chi)^2] + \mathcal{O}(\chi^3)$$
$$= \frac{1}{2} \int d^4x \sum_{n=0}^{\infty} [\dot{c}_n^2 - \omega_n^2 c_n^2] + \text{interaction}$$

Energy of each mode

$$\varepsilon = \sum_{n=0}^{\infty} \varepsilon_n = \sum_{n=0}^{\infty} \frac{1}{2} [\dot{c}_n^2(t) + \omega_n^2 c_n^2(t)]$$

As time goes, the energy is transferred to higher modes

### Universality

- Turbulence in quark mass quench
  - Time-dependent quark mass:  $m(t) = m + \delta m f(t)$



The energy spectrum approaches  $\varepsilon_n \propto \omega_n^{-5}$  as time goes

# Features of the probe brane turbulence

- The worldvolume gauge field or the quench induced by the electric field is NOT essential for the turbulence
  - We can observe the turbulence for other triggers of excitations
- Non-linearity is important
  - The initial energy spectrum excited by the quenches can be well described in linear theory
  - Even if the input is perturbative, the non-linearity will cause energy cascade from lower modes to higher modes in the time evolutions
- Wave turbulence in non-linear wave

## Summary 1

- We studied time-dependent phenomena in holographic QCD constructed by the D3/D7 system
  - We numerically solved EOMs of the probe D7-brane under time-dependent boundary conditions
- We found non-equilibrium deconfinement transition even for sub-Schwinger electric fields
  - Energy cascade from lower meson modes to higher ones
  - "turbulent meson condensation"
- We found universalities of the turbulence on the brane
  - The turbulence can occur for various excitations on the brane
  - Final energy spectrum for meson modes obeys a power law

### APPLICATION2: ROTATING ELECTRIC FIELD

### Time periodic phenomena

 Recently, time periodic states have been of great interest in condensed matter physics

- Floquet states

- Naively, holographic realization is not so easy
  - At least, two directions are inhomogeneous; EoMs are hyperbolic PDE

## Rotating electric field

- Phase transition by AC electric fields
- Time dependent system
  - Floquet states were investigated in CMP Oka, Aoki (2009)
- We can reduce EoMs to ODE in a holographic setup  $\sum E_x = E_0 \cos \Omega t$

Rotating electric field at the boundary:

$$A_x + iA_y = iE_0e^{i\Omega t}$$

Worldvolume gauge field on D7:

$$a_x(t,\rho) + ia_y(t,\rho) = b(\rho)e^{i(\Omega t - \chi(\rho))}$$

Inspired by Oka-san



### Setup

Hashimoto, SK, Murata, Oka JHEP 05 (2017) 127 SK, Murata, Oka JHEP 06 (2018) 096

DBI-action

The action becomes time independent!

$$\propto -\int d\rho \frac{\rho^3}{(w^2 + \rho^2)} \sqrt{[(\rho^2 + w^2)^2 - \Omega^2 b^2](1 + w'^2 + b'^2) + (\rho^2 + w^2)^2 b^2 \chi'^2}$$
  
static 
$$\int d\rho \frac{\rho^3}{(w^2 + \rho^2)} \sqrt{[(\rho^2 + w^2)^2 - E^2](1 + w'^2) + (\rho^2 + w^2)^2 a'^2}$$

Embedding function:  $w(\rho)$ 

 $S_{\rm D7} = -\int d^8\sigma \sqrt{-\det(h_{ab} + F_{ab})}$ 

Worldvolume gauge field:  $a_x + ia_y = b(\rho) \exp i(\Omega t + \chi(\rho))$ 

Boundary conditions at the AdS boundary

$$b|_{\rho=\infty} = \frac{E}{\Omega}, \quad \vec{E} = E \begin{pmatrix} \cos \Omega(t-t_0) \\ \sin \Omega(t-t_0) \end{pmatrix}$$
 Rotating electric field

$$w(\rho) = m + \frac{c}{\rho^2} + \cdots, \quad \vec{a}(t,\rho) = -\int^t dt' \vec{E}(t') + \frac{1}{2\rho^2} \vec{j}(t) + \cdots$$

### Symmetry

### Static electric field

- Time translation: yes
- Rotation: no
- Helical: no

### Rotating electric field

- Time translation: no
- Rotation: no
- Helical: yes



Time translation in static cases changes to helical (corotating) translation in rotating cases

• EoM  

$$b'' = \frac{1}{\rho(w^2 + \rho^2)} \frac{1}{(w^2 + \rho^2)^2 - \Omega^2 b^2} \left\{ -3(w^2 + \rho^2)^3 b'(1 + w'^2 + b'^2) - \Omega^2 b(1 + w'^2 + b'^2) \left[ \rho(w^2 + \rho^2) - 0b'(3w^2 + \rho^2) \right] + b(w^2 + \rho^2)^3 (\rho - 3bb')\chi'^2 \right\},$$

$$\chi'' = \frac{1}{b\rho(w^2 + \rho^2)} \frac{1}{(w^2 + \rho^2)^2 - \Omega^2 b^2} \left\{ -3b^3 \left( \frac{2}{\rho} - 233 - t^3 \right) \right\}$$
Denominator will vanish  $\Rightarrow$   
 $- \left[ 2\rho b'(w^2 + \rho^2)^3 + 3b(w^2 + \rho^2)^3 (1 + w'^2 + b'^2) \right] \chi' \right\},$ 

$$w'' = \frac{1}{\rho(w^2 + \rho^2)} \frac{1}{(w^2 + \rho^2)^2 - \Omega^2 b^2} \left\{ -3(w^2 + \rho^2)^3 (1 + w'^2 + b'^2)w' + \Omega^2 b^2 \left[ 2\rho w + (\rho^2 + 3w^2)w' \right] (1 + w'^2 + b'^2) - 3b^2(w^2 + \rho^2)^3 w'\chi'^2 \right\}.$$

We can obtain solutions in the same way as in static cases

### Phase diagram



- As the frequency increases, the critical value becomes smaller
- Near frequencies of the meson masses, the critical value is infinitesimal

### Phase diagram

Fine structure around the lowest mode



- Multiple phases of insulator and conductor states coexist in ٠ vicinity of the critical curve Even in infinitesimal electric fields, there are non-trivial states
- other than the vacuum state

## Summary 2

- The rotating electric fields can provide a new axis of the phase diagram
  - AC Schwinger limit is smoothly connected with DC Schwinger limit in static electric field
- Frequencies of the rotating field can reduce the critical amplitude to cause dielectric breakdown
  - The AC Schwinger limit periodically becomes infinitesimal near the frequencies corresponding to the meson mass spectrum
- Even in E = 0, non-trivial states emerge other than the vacuum state
  - Non-linear excitation of vector meson, Floquet condensation of vector meson



## Sketch from bulk viewpoint



# Revisit of deconfinement and thermalization

- If static, both of deconfinement and thermalization are given by the same condition in gravity side: the horizon exists or not
  - Quasi-normal modes = unstable mesons
  - Hawking temperature = thermalization
- In time-dependent cases, this definition is not so useful
  - When does the horizon form for temporal observers at the AdS boundary?
    - By definition, "horizon" is invisible from the AdS boundary (=null infinity) forever.
  - Since we have no preferred time-slice in gravity, the formation time is ambiguous

# Redshift factor and surface gravity

- Redshift factor
  - The ratio between the energy observed on the AdS boundary and the initial surface (static region)

$$R(u_0) = \frac{k^a \xi_a|_{v=v_{\text{ini}}}}{k^a \xi_a|_{v=u_0}} = -\frac{m}{2} \frac{V_{,v}(u_0, u_0)}{\Phi_{,u}(u_0, v_{\text{ini}})}$$

- Surface gravity
  - Relation between times to define the Initial state and the final state  $\kappa(u_0) = \frac{d}{du} \log R(u_0)$

AdS boundary

If this quantity is almost constant, it becomes Hawking temperature observed at  $u_0$ 

Redshift becomes too large  $\Leftrightarrow$  deconfinment Surface gravity becomes constant  $\Leftrightarrow$  thermalization

We can define and calculate these only from the causal past of temporal observers

## Anti-de Sitter(AdS) space

 Vacuum solution of the Einstein eq. with negative cosmological constant

